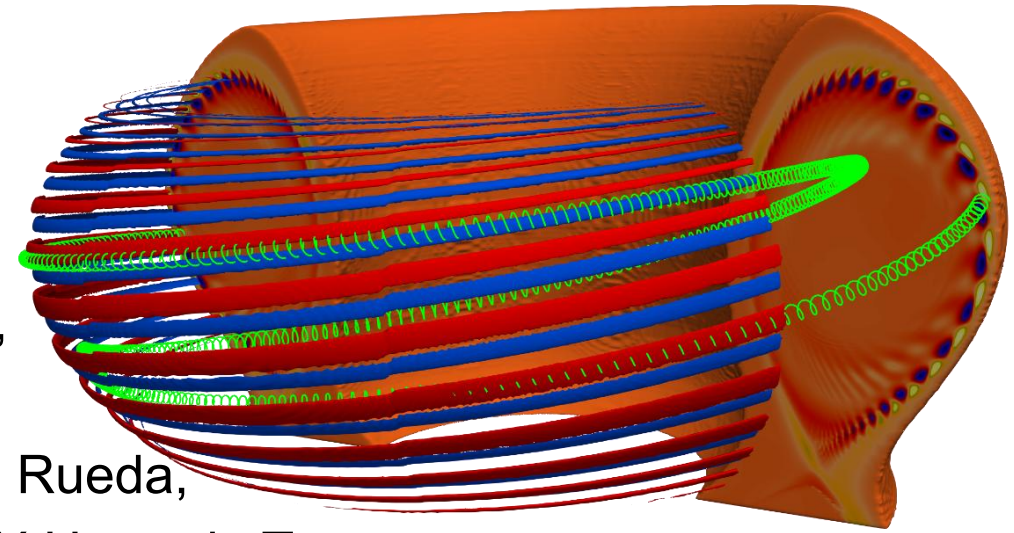


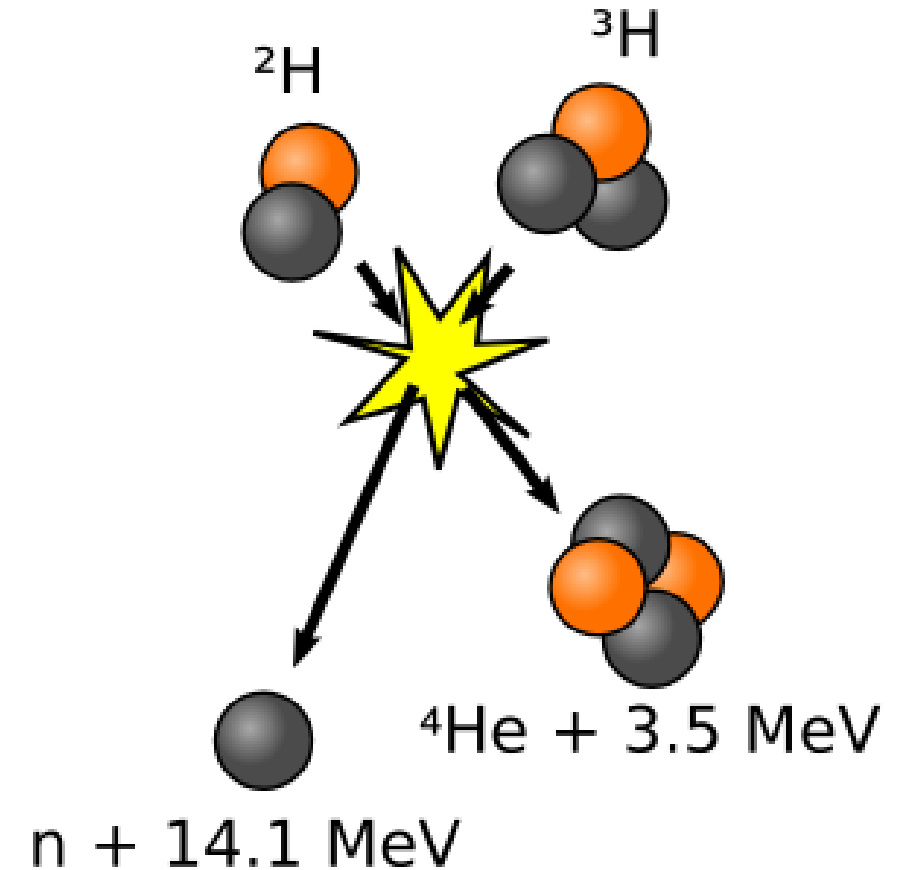
3D hybrid kinetic-MHD modelling of the interaction between Edge Localised Modes and Energetic Particles in the ASDEX Upgrade tokamak

J.J. Domínguez-Palacios Durán, M. García Muñoz, E. Viezzer, S. Futatani, M. Toscano Jiménez, J. Galdón Quiroga, J. González Martín, A. Jansen van Vuuren, P. Oyola, J. Rueda Rueda, L. Sanchís, Y. Suzuki, Y. Todo, EUROfusion WPTE and ASDEX Upgrade Teams



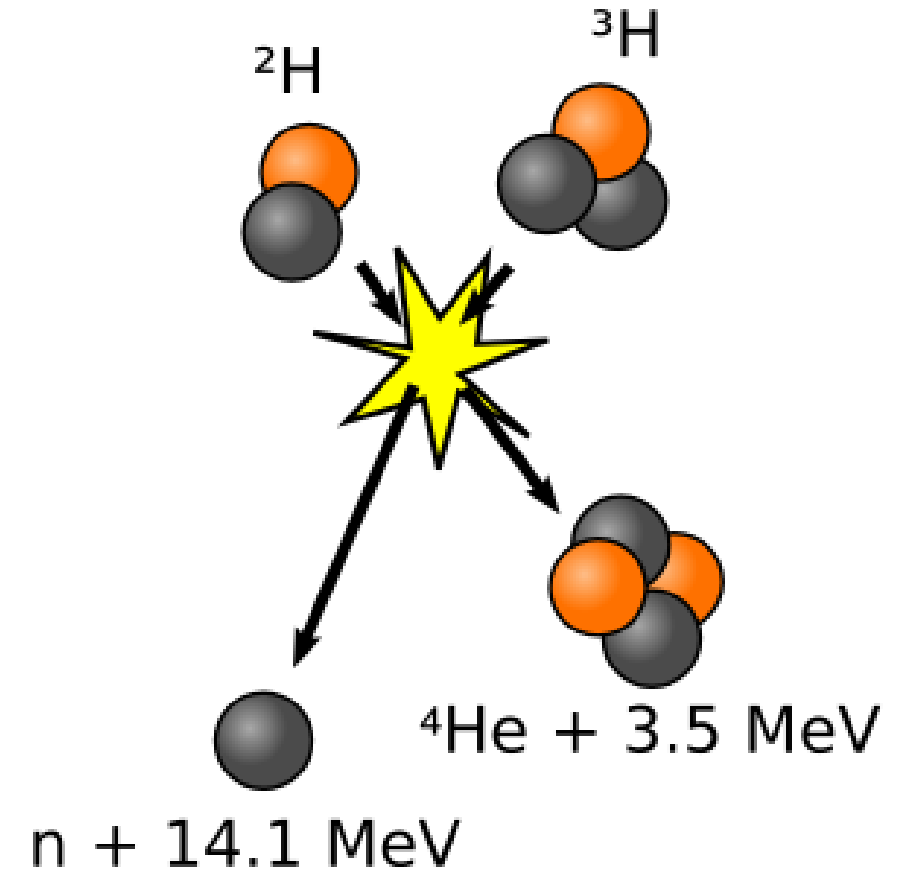
Nuclear fusion is a promising energy source

- Mass deficit leads to large energy release due to $E = mc^2$



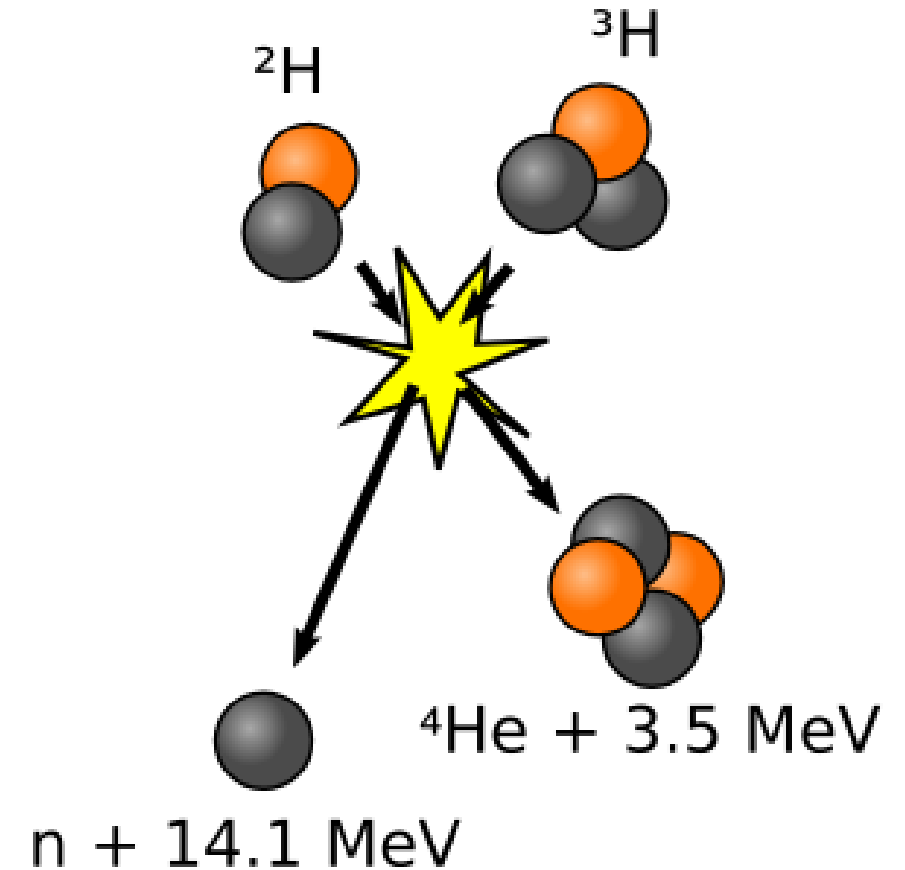
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- 17.6 MeV is a huge number...



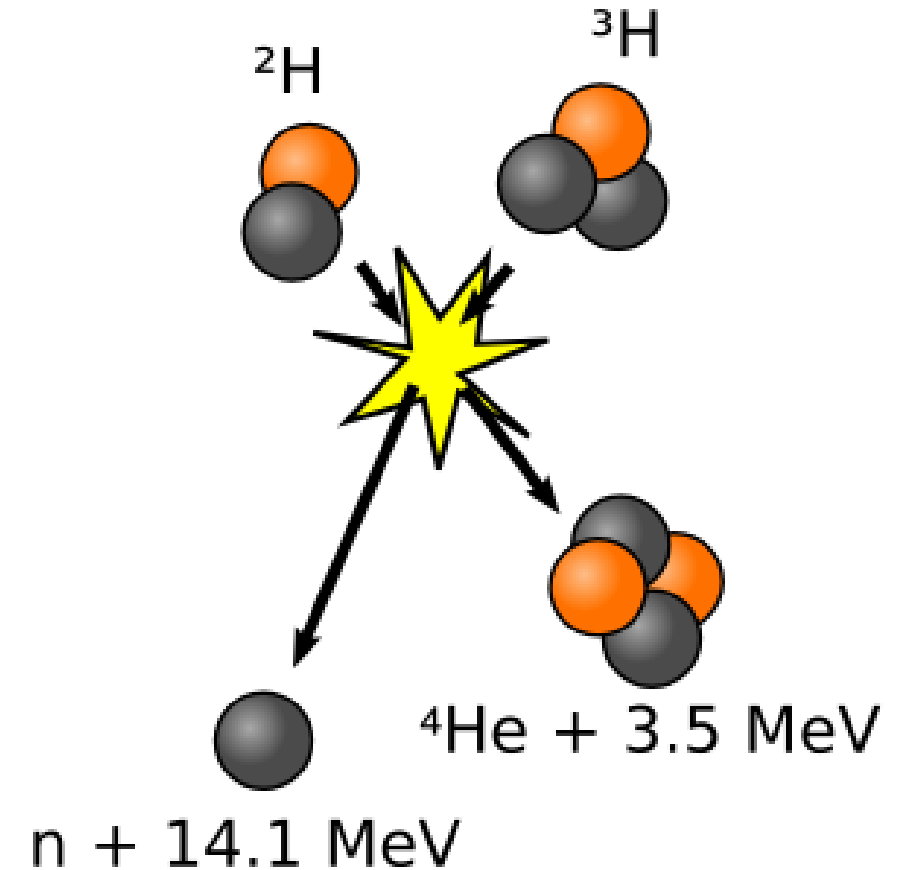
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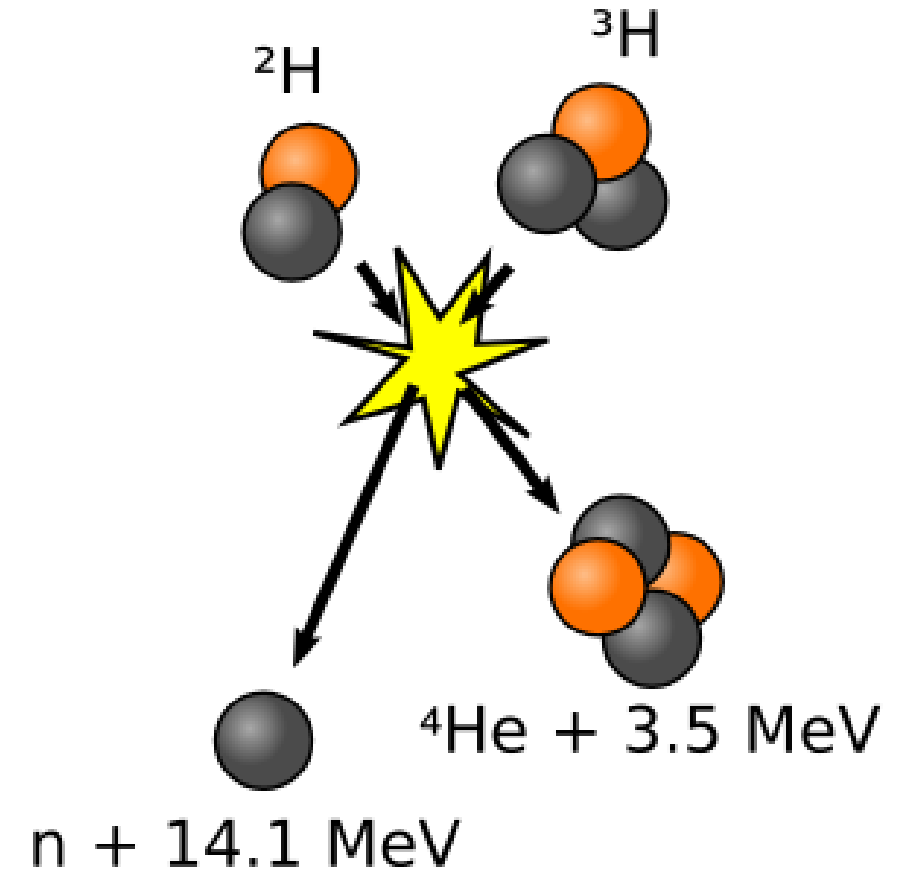
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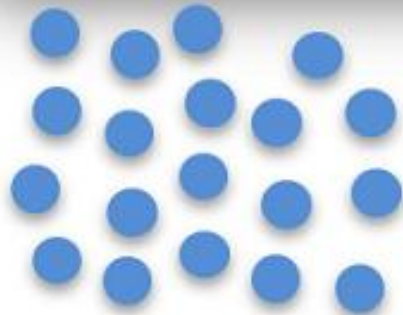
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 - Fission of ^{235}U releases $\sim 10^8$ kJ/g
 - D-T fusion releases $\sim 10^{11}$ kJ/g



The fusion fuel is, and will be, in the plasma state in present and future fusion devices

Gas

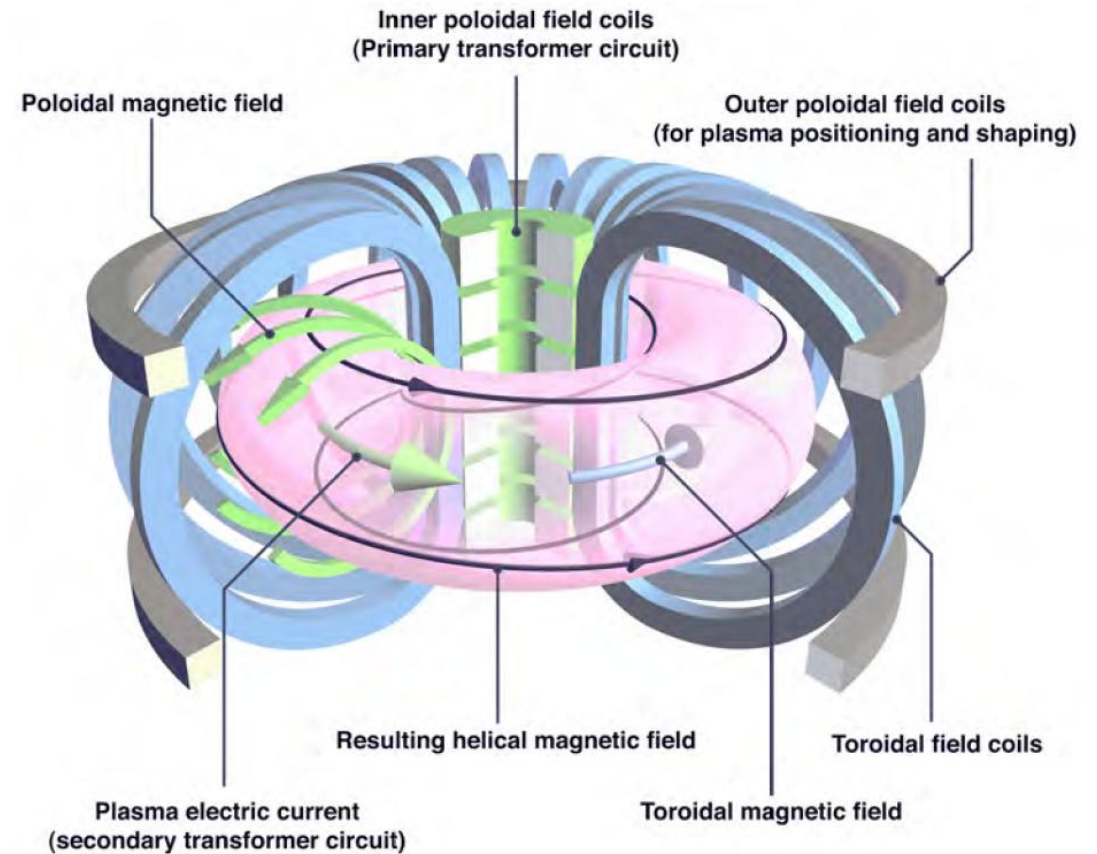


Plasma



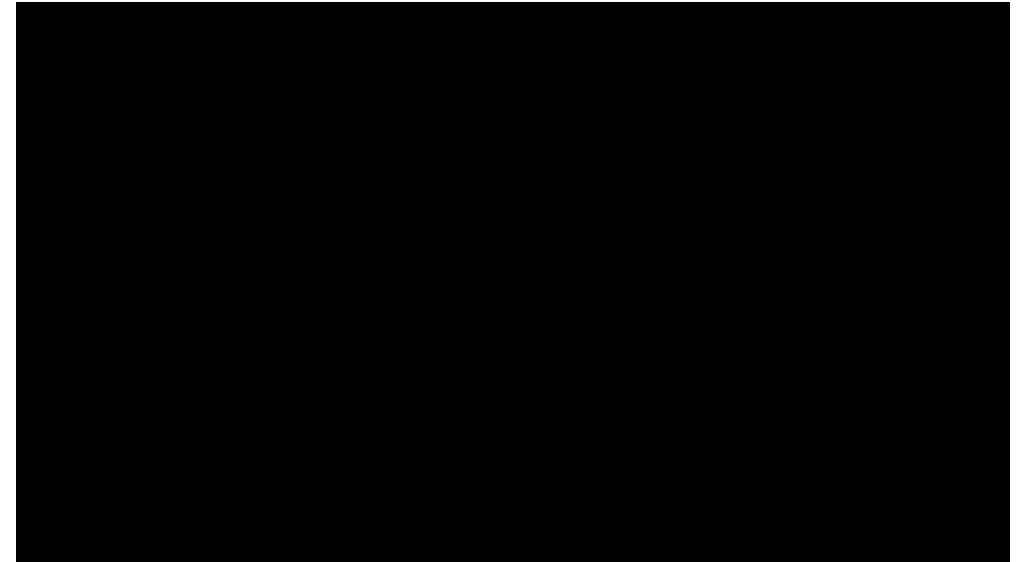
The tokamak concept

- Magnetic confinement fusion device with torus shape
- Toroidal + poloidal magnetic field
- Poloidal field coils control plasma column position and shape of poloidal cross section



Future tokamaks are expected to operate in High confinement mode (H-mode)

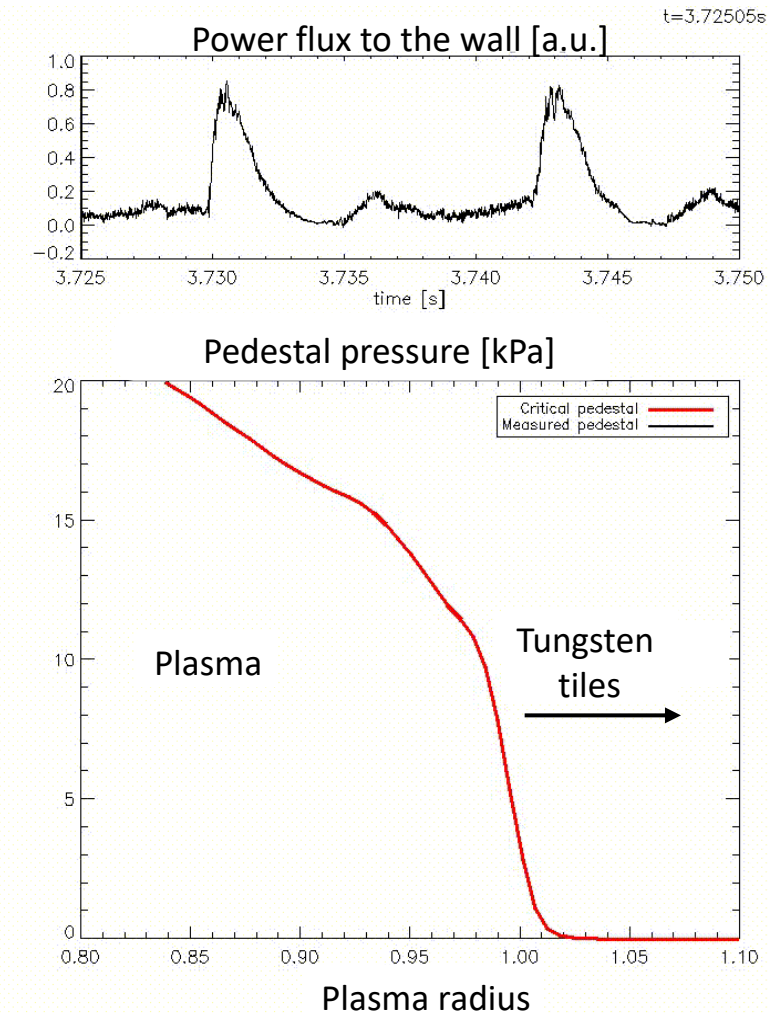
- H-mode has shown highest confinement rates in tokamaks



New Fusion Energy Record Pulse from JET DTE3 campaign - Pulse #104522
69 MJ in 6 second-long pulse fusing 0.2 mg of fusion fuel
<https://www.youtube.com/watch?v=hQVzroA9qsE>

Future tokamaks are expected to operate in High confinement mode (H-mode)

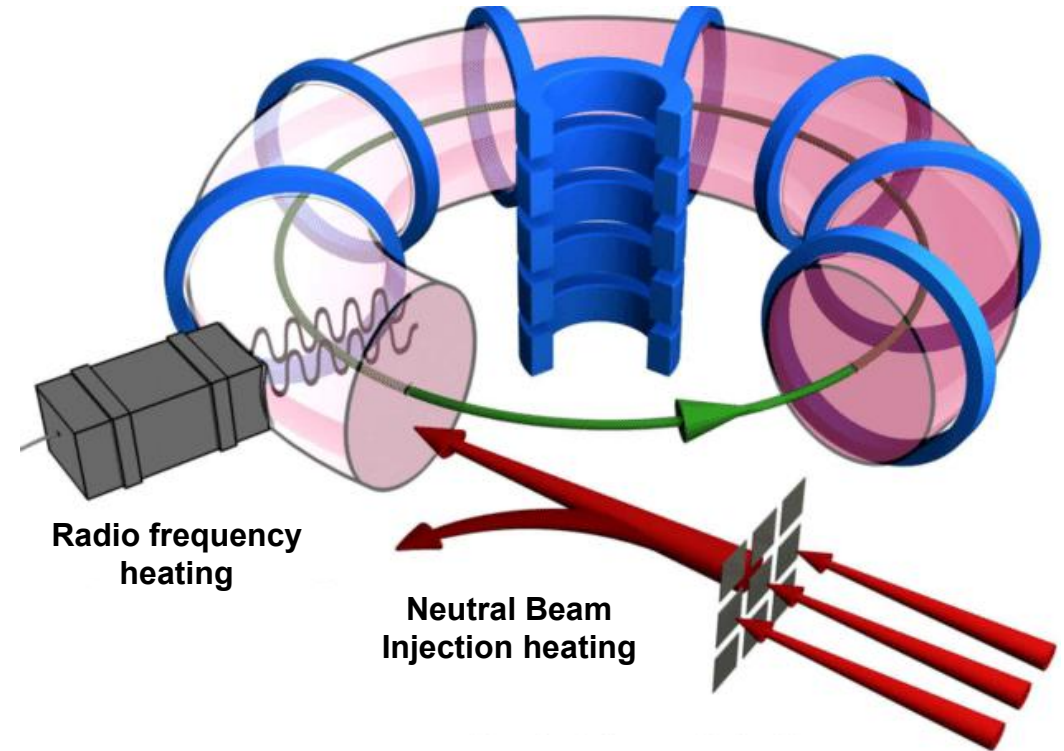
- H-mode has shown highest confinement rates in tokamaks
- Limited by ELMs^[1]
 - Degradation of pedestal
 - Intolerable for future fusion devices



*A. Burckhart

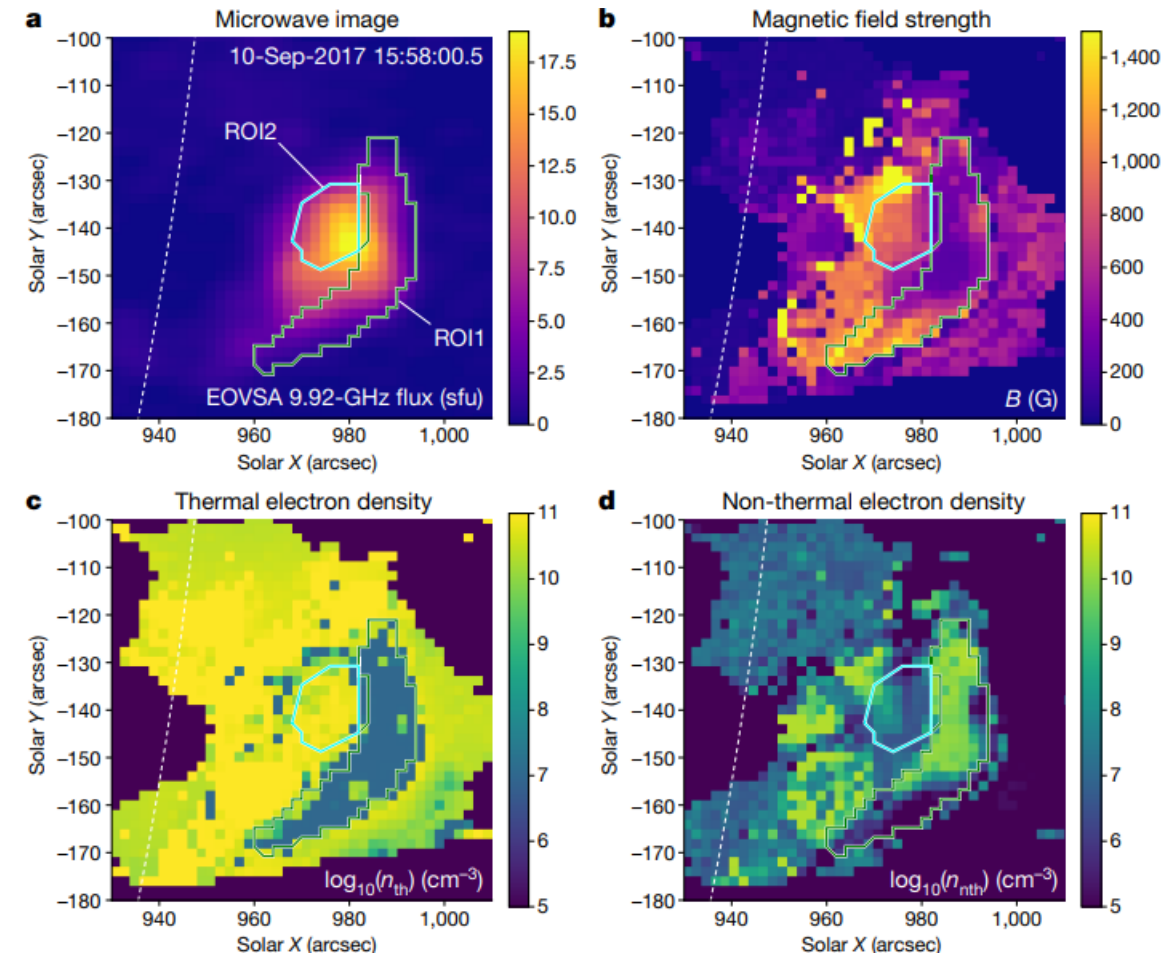
Energetic particles are used to heat up the plasma

- Main heating mechanisms
 - ICRF, ECRF (external)
 - NBI (external)
 - Fusion-born product (internal)



Energetic particles are used to heat up the plasma

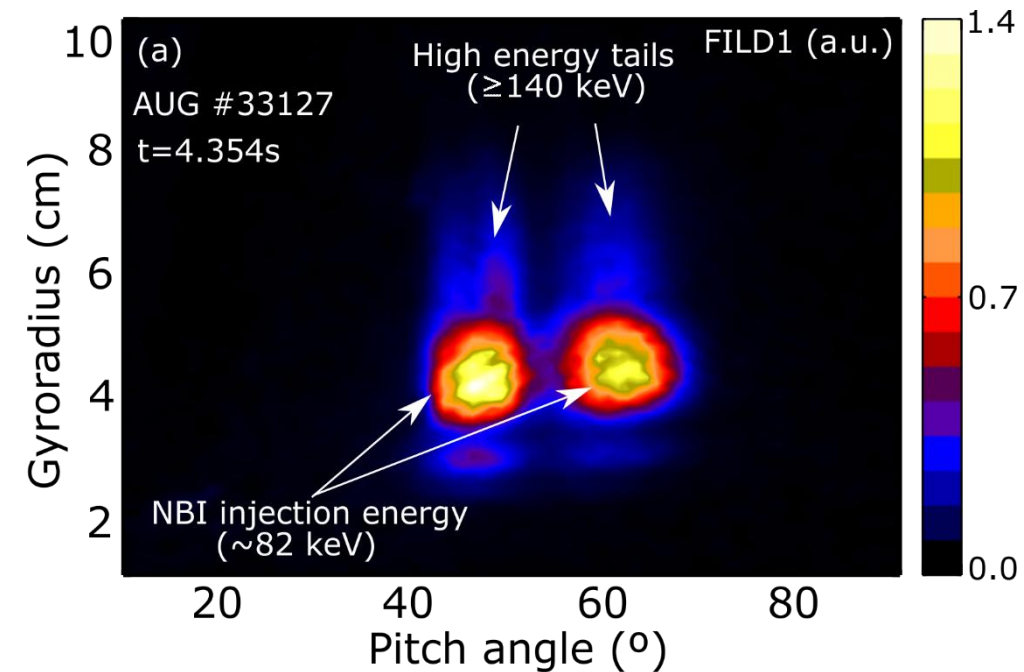
- Main heating mechanisms
 - ICRF, ECRF (external)
 - NBI (external)
 - Fusion-born product (internal)
- Energetic particles are prone to a rich variety of wave-particle interactions due to their large velocities and long mean free path



G.D. Fleishman *et al.*, Nature **606**, 674 (2022)

Experimental observations suggest interaction between ELMs and fast-ions

- Beam-ion acceleration during ELM crashes
- **Fast-ion kinetic effects on ELMs are needed!**



*J. Galdon-Quiroga *et al.*, 2018, Phys. Rev. Lett. **121**, 025002

- Hybrid kinetic-MHD MEGA simulation code
- Simulation results
- Conclusions & Outlook

- Hybrid kinetic-MHD MEGA simulation code
- Simulation results
- Conclusions & Outlook

Navier-Stokes equations (1823, 1845) describe fluid dynamics



Navier
(Wikipedia)



Stokes
(Wikipedia)

$$\rho \frac{d\vec{u}}{dt} = -\nabla p + \rho \vec{g} - \nabla \cdot \vec{\Pi}$$

C.L. Navier, Sur les Loix du Mouvement des Fluides, Mémoires de L'Académie Royale des Sciences de L'Institut de France, Paris, VI, 389-440 (1827)
G. G. Stokes, On the Theories of the Internal Friction of Fluids in Motion, and of the Equilibrium and Motion of Elastic Solids. Cambridge University Press, 1880, pp. 75-129

MHD equations: the plasma is regarded as an electromagnetic fluid (Alfvén, 1942)



Navier
(Wikipedia)



Stokes
(Wikipedia)



Alfvén
(ITER website)

$$\rho \frac{d\vec{u}}{dt} = -\nabla p + \vec{j} \times \vec{B} - \nabla \cdot \vec{\Pi}$$

MEGA: hybrid kinetic-MHD code to study the interaction between thermal plasma and energetic particles



Navier
(Wikipedia)



Stokes
(Wikipedia)



Alfvén
(ITER website)



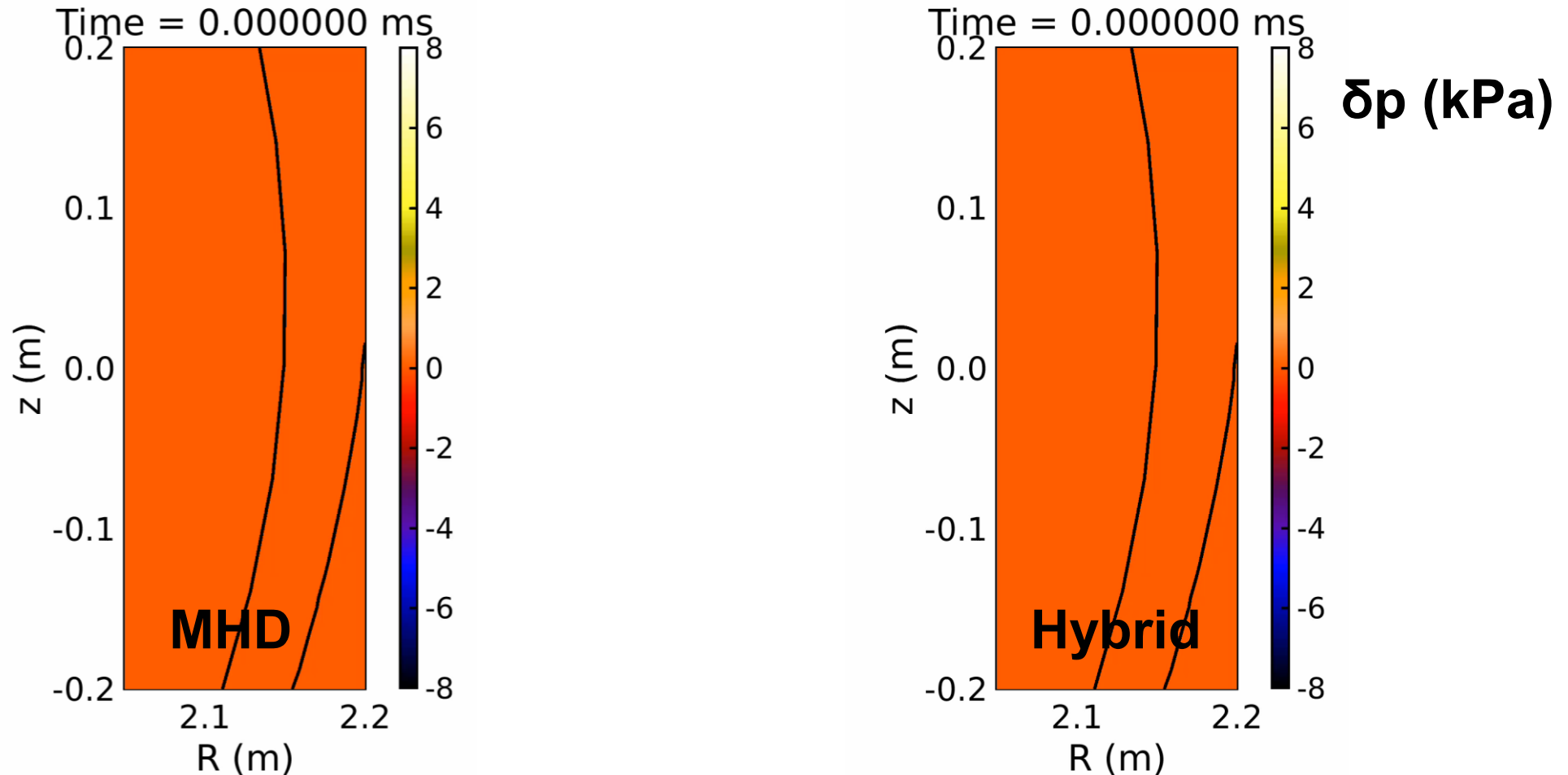
Prof. Yasushi Todo
(NIFS website)

$$\rho \frac{d\vec{u}}{dt} = -\nabla p + (\vec{j} - \vec{j}_h) \times \vec{B} - \nabla \cdot \vec{\Pi}$$

MEGA calculates the self-consistent interaction between **MHD** and **EPs**

- Hybrid kinetic-MHD MEGA simulation code
- Simulation results
- Conclusions & Outlook

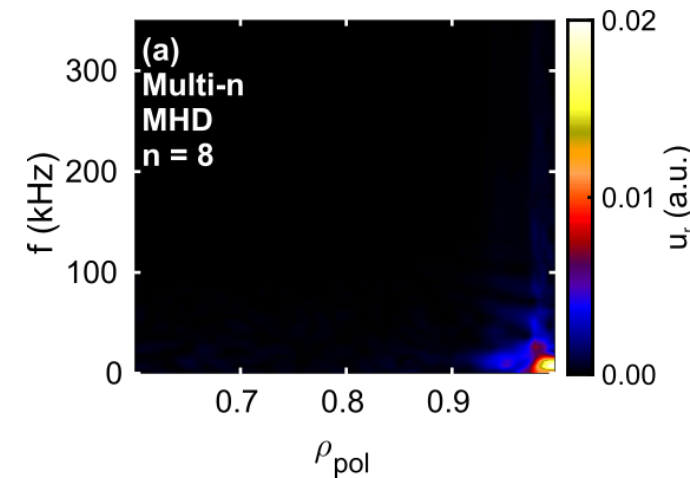
Energetic-ions significantly modify perturbation structure



- The interaction with fast-ions modify the plasma flow pattern and, consequently, the shear of the structure

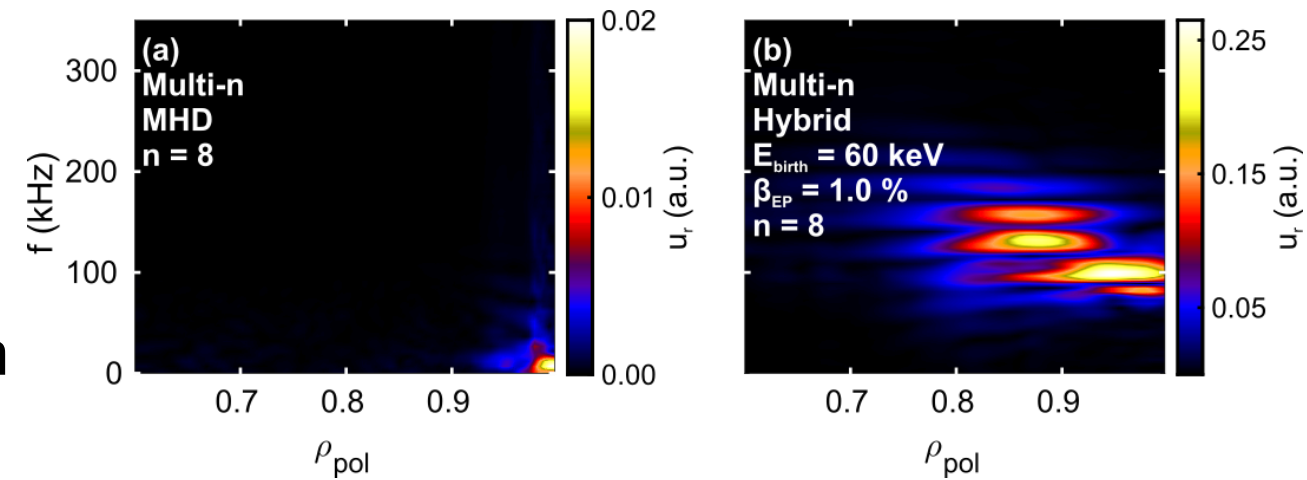
Spatial-frequency pattern of modes during ELM crash changes abruptly in the presence of fast-ions

- In pure MHD simulations, $f \sim 20$ kHz



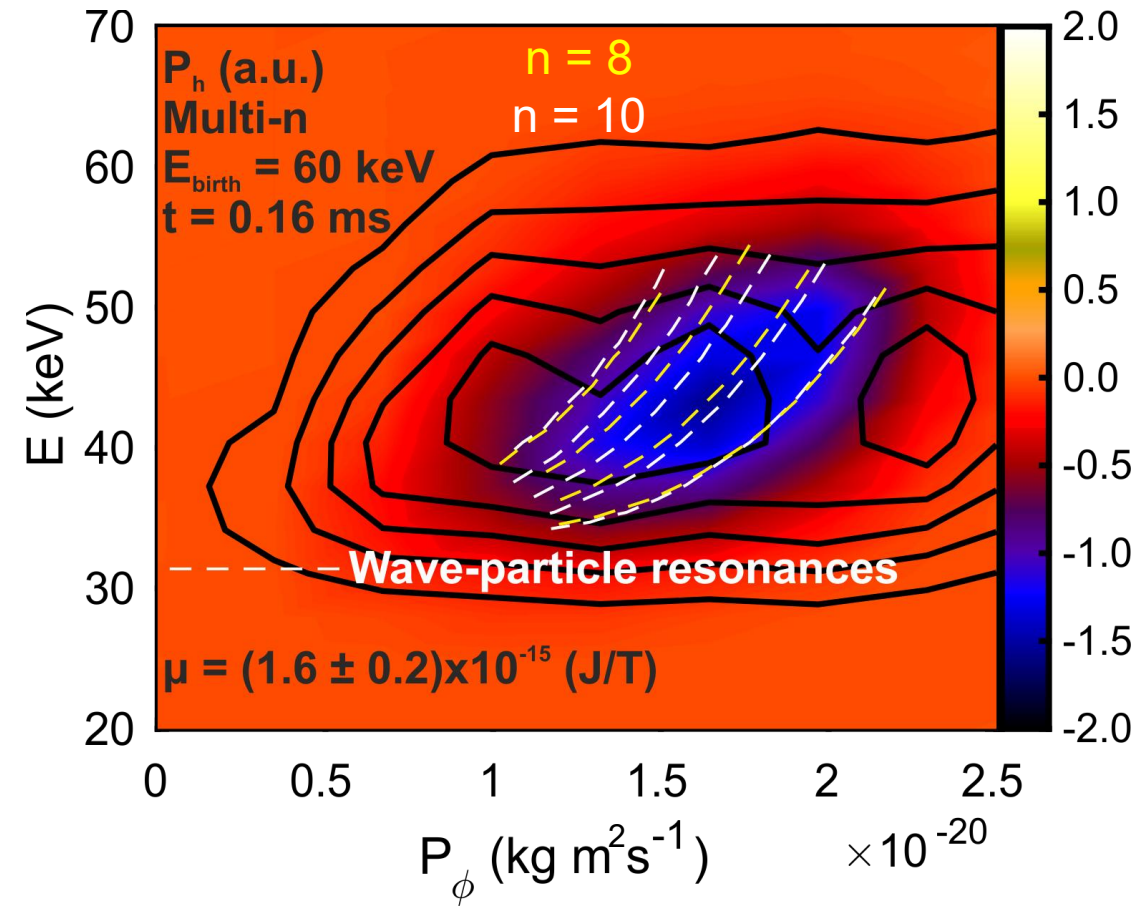
Spatial-frequency pattern of modes during ELM crash changes abruptly in the presence of fast-ions

- In pure MHD simulations, $f \sim 20$ kHz
- In hybrid multi-n simulations, $f \sim 200$ kHz ($n = 8$)
- Modes also become radially broad in the presence of fast-ions

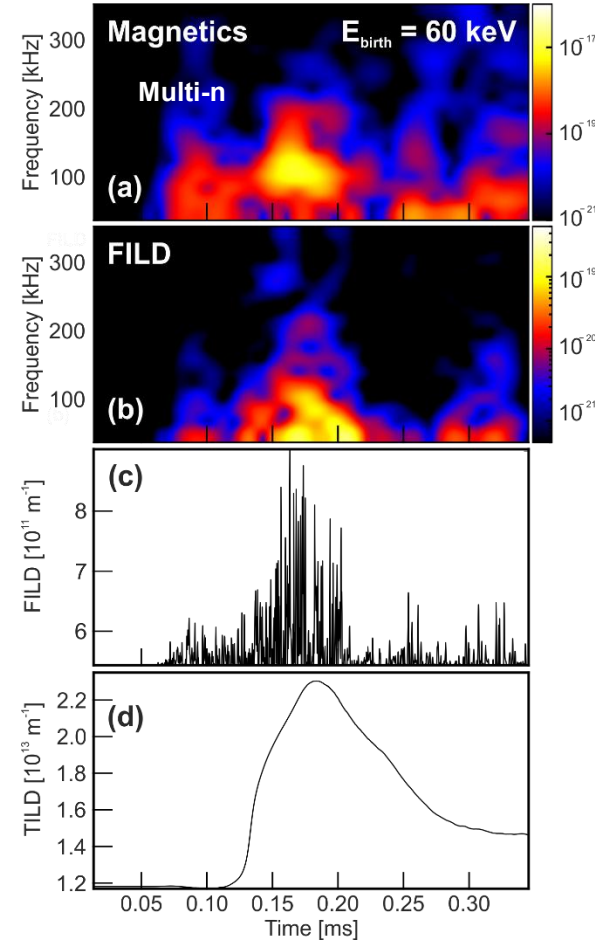
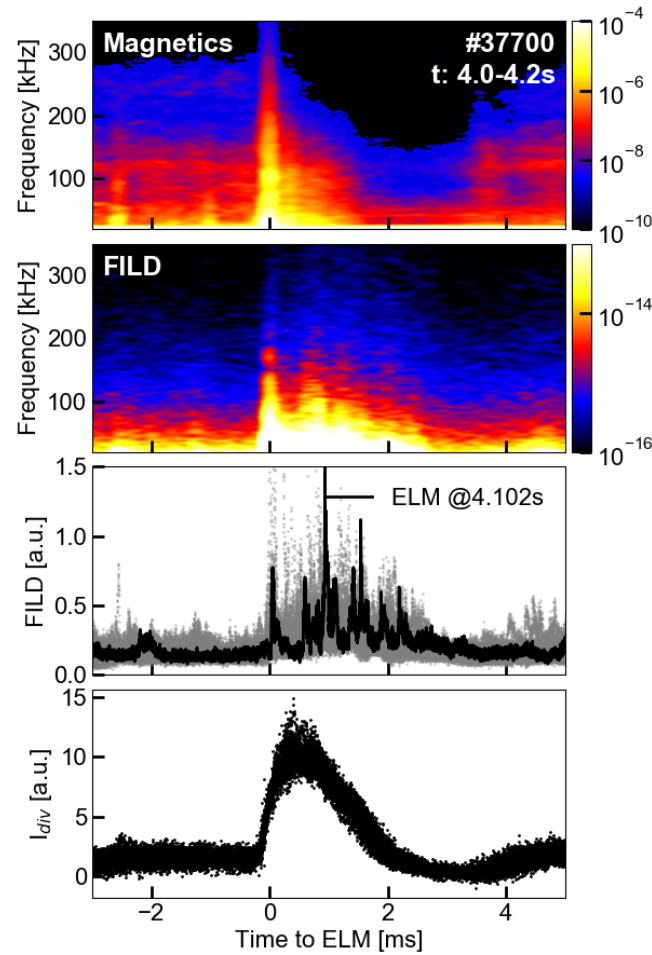


Role of energetic particle force is due to a resonant interaction with the electromagnetic perturbation

- **Power transfer:** $P_h = \sum_l w_l \frac{dE_l}{dt}$
- **Resonance:** $\omega_n - n\omega_\phi - p\omega_\theta \approx 0$
- Resonant interaction leads to net energy exchange that modifies ELM



MEGA synthetic diagnostics reproduce some key experimental observations



- Simulations qualitatively reproduce abrupt crash with magnetic and FILD frequency patterns extended to high frequencies. Divertor thermal currents show similar behaviour

- Motivation
- Hybrid kinetic-MHD MEGA simulation code
- Simulation results
- Conclusions & Outlook

- Fast-ion kinetic effects impact radial width, frequency and shear of perturbation
- Interaction mechanism: **resonant energy exchange between electromagnetic field of ELM and fast-ions**
- Fast-ion kinetic effects qualitatively help to understand experimental observations

Some open questions derived from this thesis

- What is the role of fast-ions in regimes absent of ELMs?
- What will happen in future burning plasmas?
 - Will fast-ions help to reduce ELM-induced heat flux on the divertor materials? Or do they have the opposite effect?

THANK YOU!

Energetic particle dynamics in MEGA (1/2)

- Drift-kinetic description of particle motion

$$\vec{u}_{gc} = \vec{v}_{\parallel}^* + \vec{v}_E^* + \vec{v}_B^*$$

$$\vec{v}_{\parallel}^* = \frac{v_{\parallel}}{B^*} (\vec{B} + \rho_{\parallel} B \nabla \times \hat{b})$$

$$\vec{v}_E^* = \frac{1}{B^*} \vec{E} \times \hat{b}, \vec{v}_B^* = -\frac{\mu}{q_h B^*} \nabla B \times \hat{b}$$

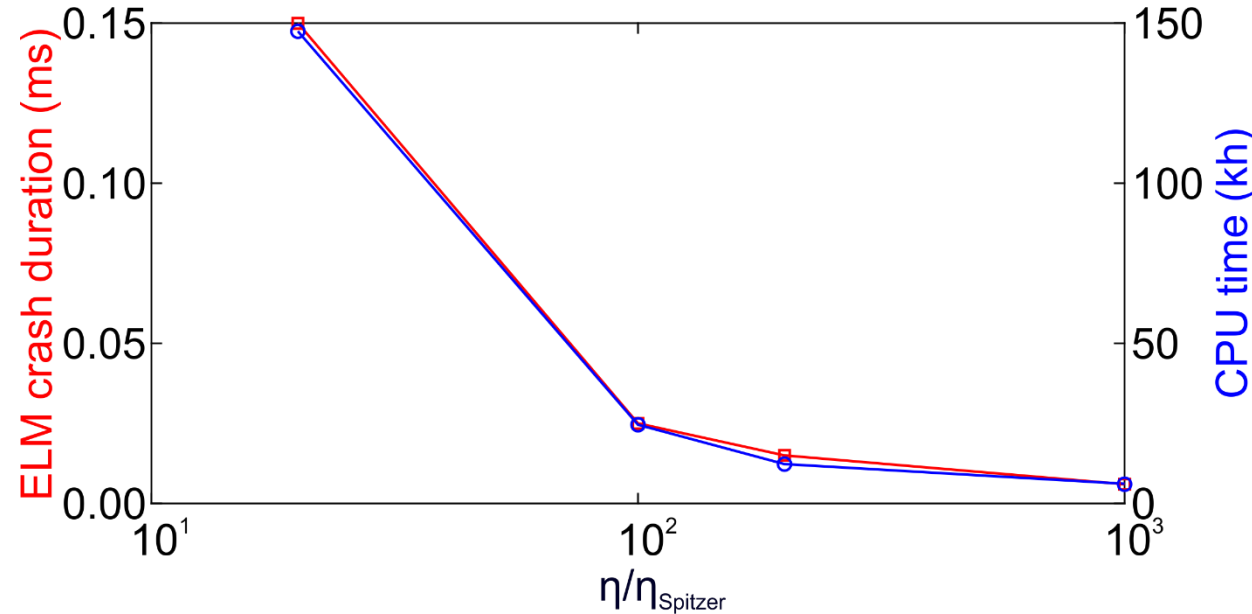
$$\rho_{\parallel} = \frac{m_h v_{\parallel}}{q_h B}, \mu = \frac{m_h v_{\perp}^2}{2B}$$

$$\vec{B}^* = B(1 + \rho_{\parallel} \hat{b} \cdot \nabla \times \hat{b})$$

$$m_h v_{\parallel} \frac{dv_{\parallel}}{dt} = \vec{v}_{\parallel}^* \cdot (q_h \vec{E} - \mu \nabla B)$$

- Electromagnetic field gyro-averaged (FLR effects)

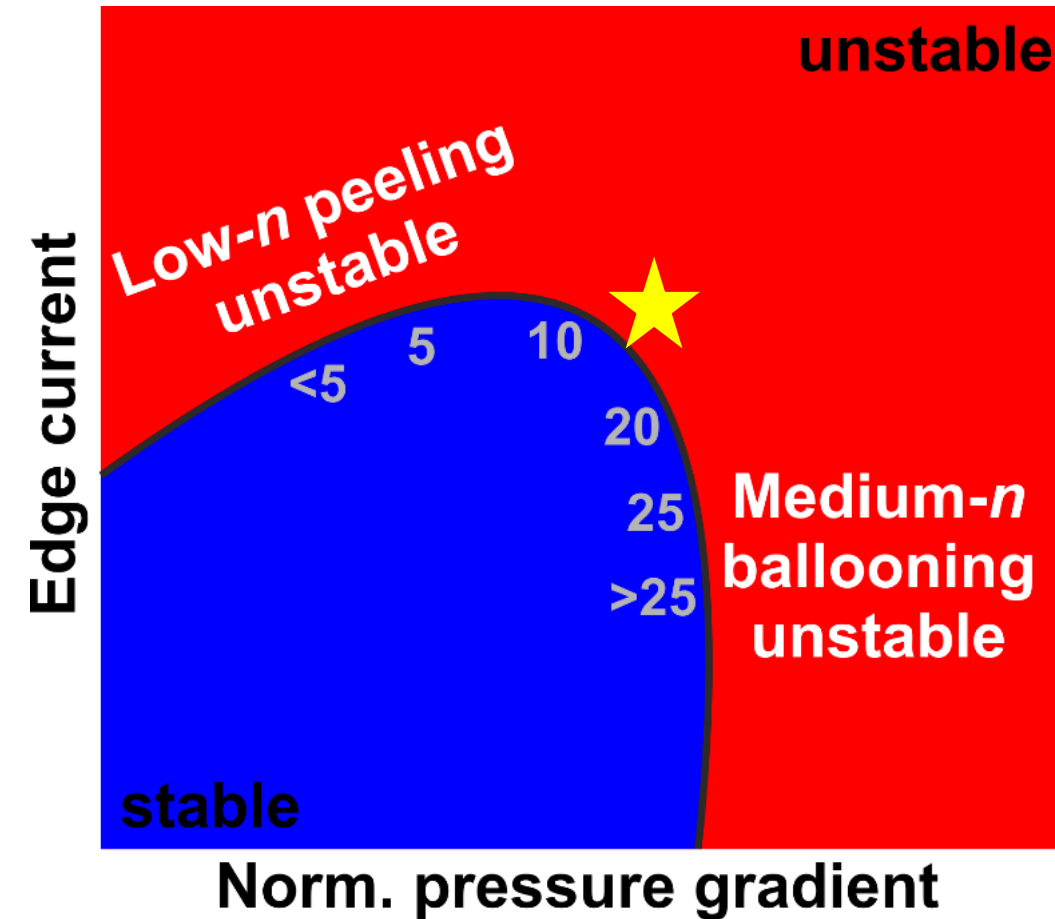
ELM crash duration strongly depends on plasma resistivity



- ELM crash duration is shorter when η decreases
- To simulate with $\eta = \eta_{\text{Spitzer}}$ and ELM crash duration of 1-2 ms, 130 years of wall-time are needed
- We expect to see same striking similarities in synthetic diagnostics when simulation with $\eta = \eta_{\text{Spitzer}}$ is affordable

How to “best” describe ELMs theoretically?

- Linear MHD: coupled peeling-ballooning modes driven by ∇p and j ^[1]
- Linear MHD is not enough because...
 - Cannot capture growing and mitigation phases of instabilities
 - Impossible to simulate heat and particle transport due to ELMs with linear MHD
- **Nonlinear models are essential to simulate the key ELM physics** ^[2]

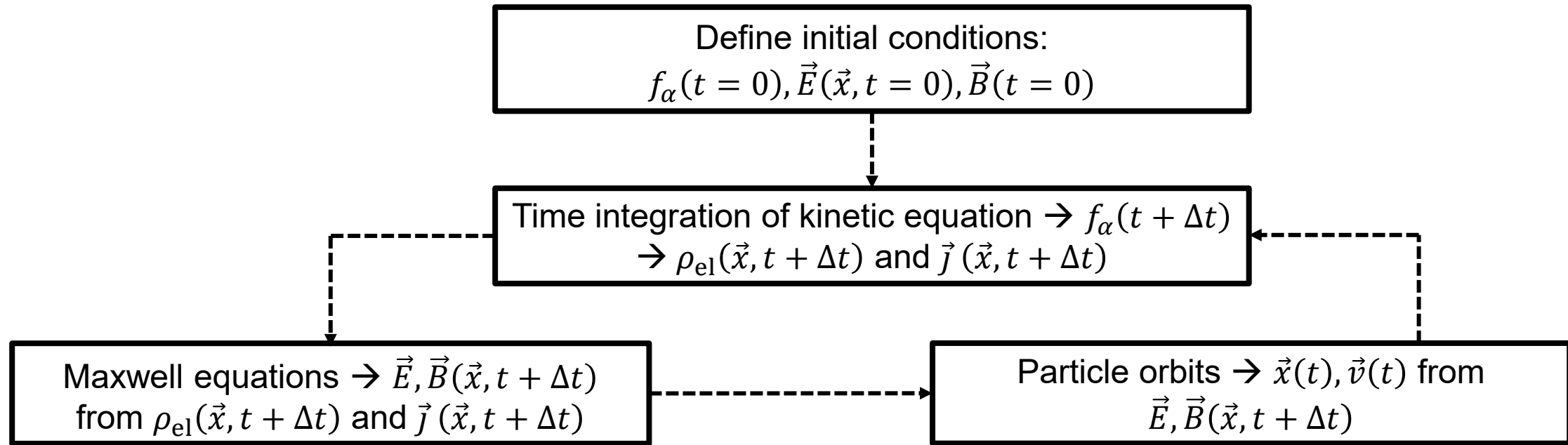


*E. Viezzer

[1] J.W. Connor *et al.*, Phys. Plasmas **5**, 2687 (1998)

[2] G.T.A. Huysmans *et al.*, Phys. Plasmas **22**, 021805 (2015)

Solving the kinetic equation



- Difficulties of kinetic equation

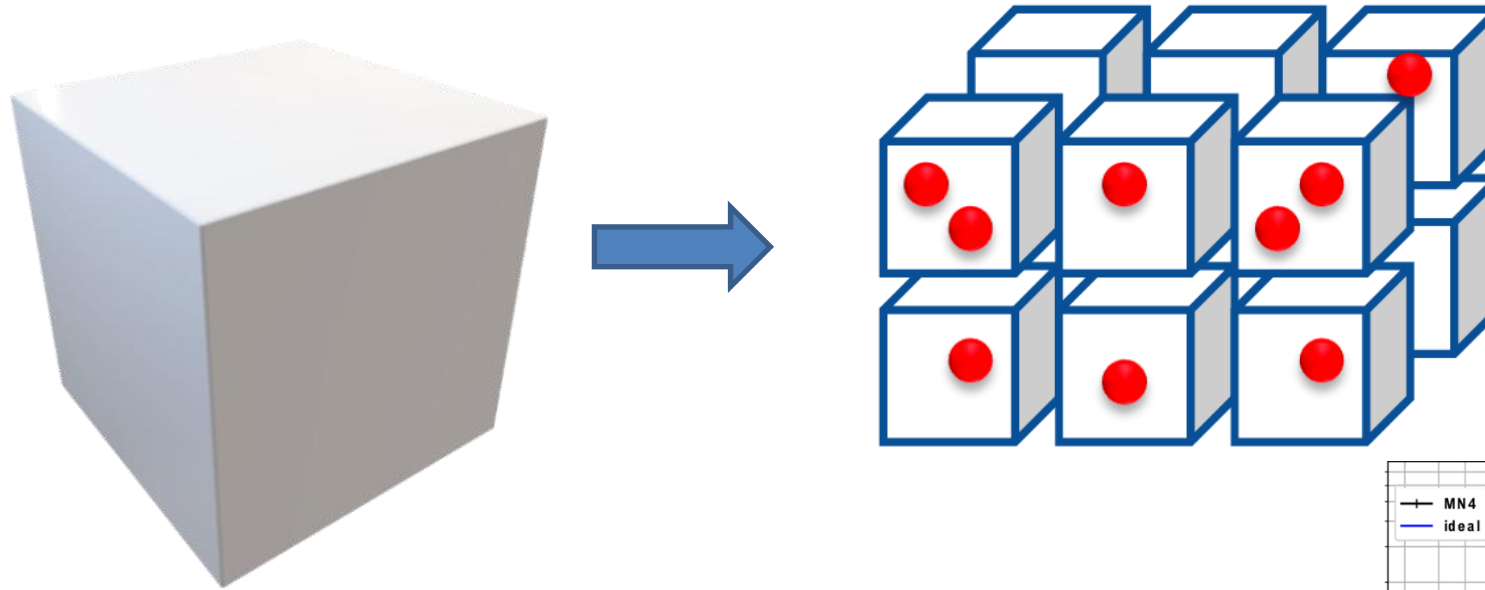
- Not possible to solve analytically for most of cases of interest
- No universal choice of collision term
- Still has a relatively large size

Why do we need a kinetic equation?

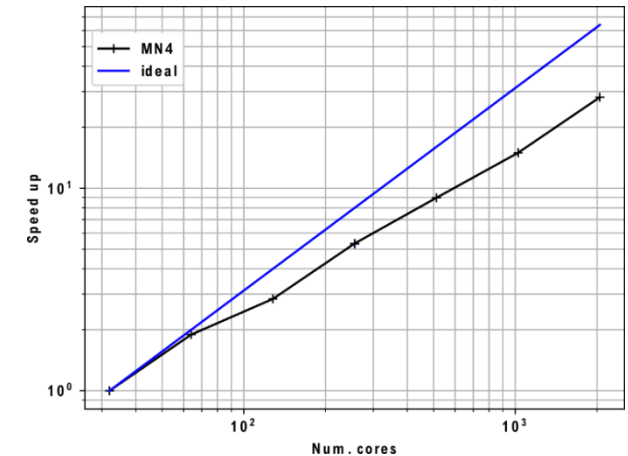
- Solving particle equation of motion is not feasible because
 - The exact form of the electromagnetic field is not known a priori → impossible to carry out an analytical study
 - There are 10^{23} particles → impossible to solve computationally such a many-body problem
- Statistical description of the plasma: **kinetic equation**

Parallelization used in the simulations

- Particles and fluid fields are 3 dimensionally sub-divided
→ MPI parallelization



- Good scaling until 2048 cores of MareNostrum IV



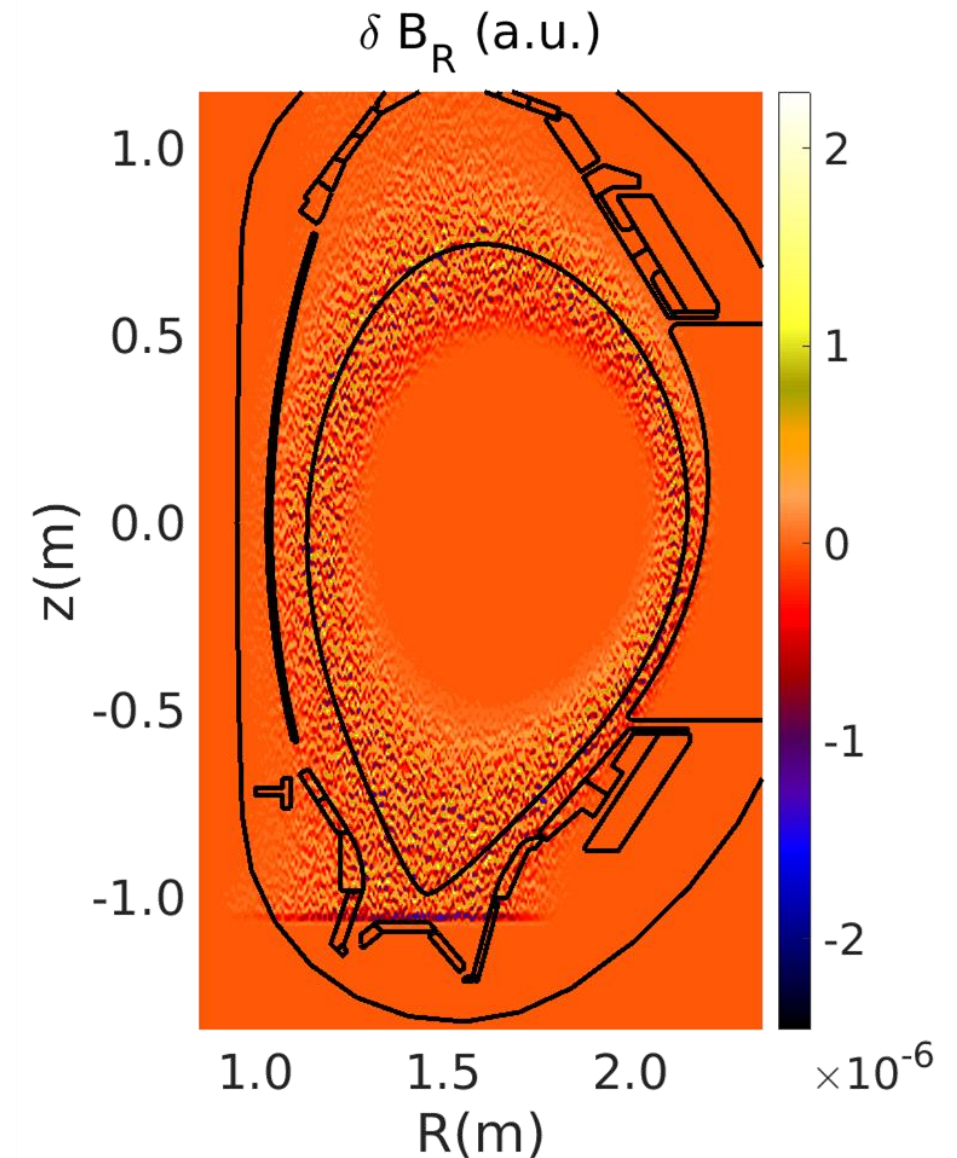
*Courtesy by J. Gonzalez-Martin

Initial seed perturbation in MEGA

- Initial magnetic field perturbation $\delta \vec{B}$ at noise level and given by

$$\delta \vec{B} = \nabla \times \vec{F}$$

- $\vec{F}$ denotes a random vector field
- Divergence-free nature of magnetic field ensured^[1]



[1] Y. Hu *et al.*, Phys. Plasmas **23**, 022505 (2016)

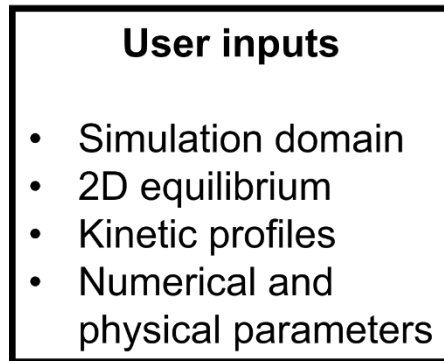
(1)

User inputs

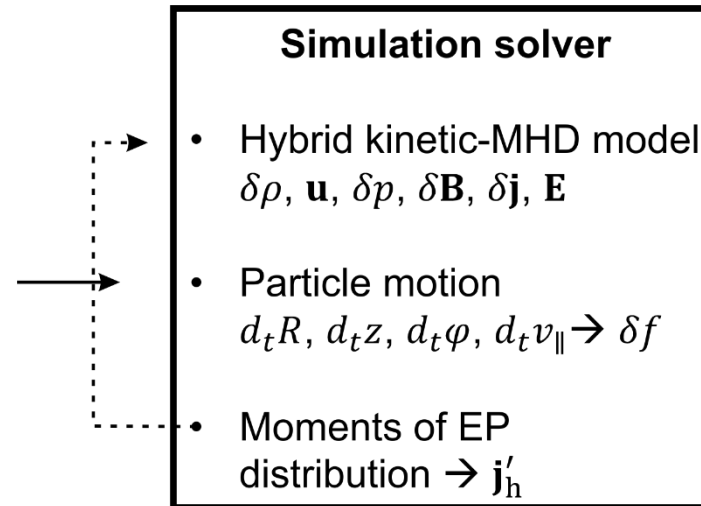
- Simulation domain
- 2D equilibrium
- Kinetic profiles
- Numerical and physical parameters

General workflow of MEGA

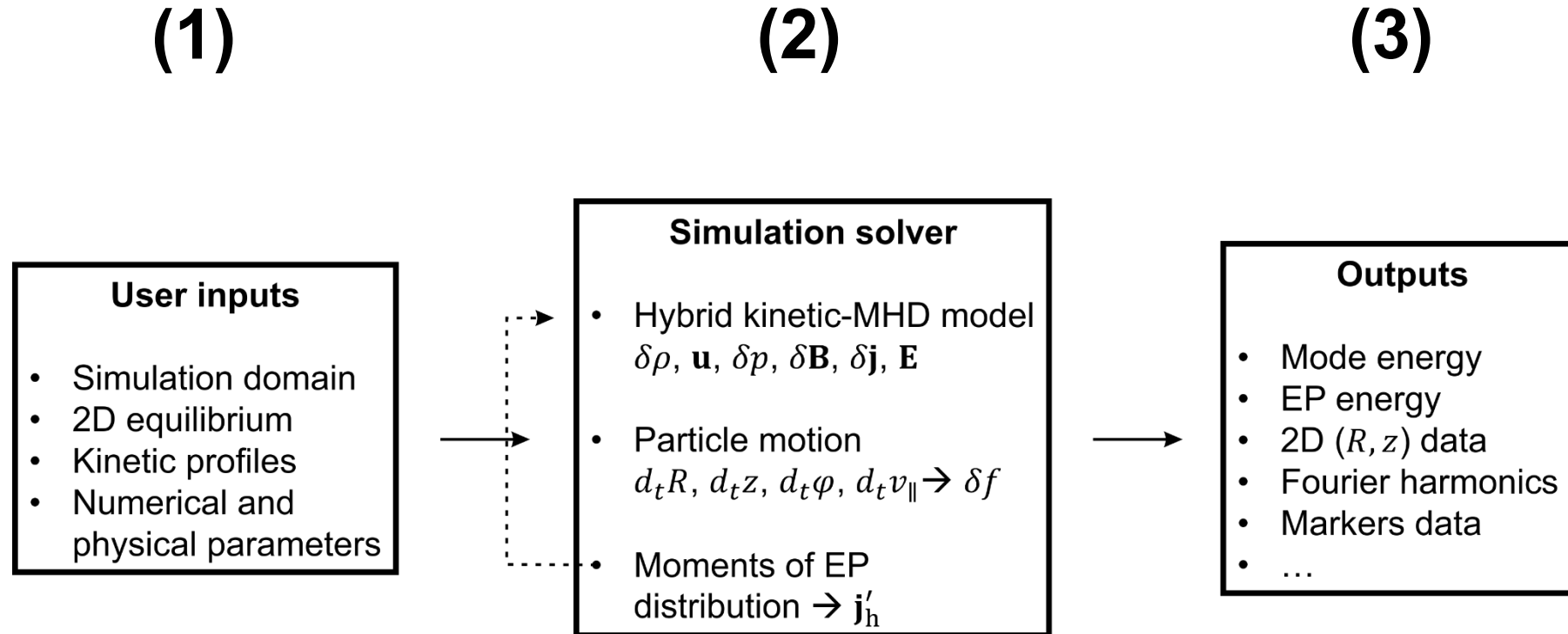
(1)



(2)



General workflow of MEGA



- Kinetic equation solved with δf method $\rightarrow f = f_0 + \delta f$
- For collisionless particles, noting that $\vec{z} \equiv (\vec{x}, \vec{v})$

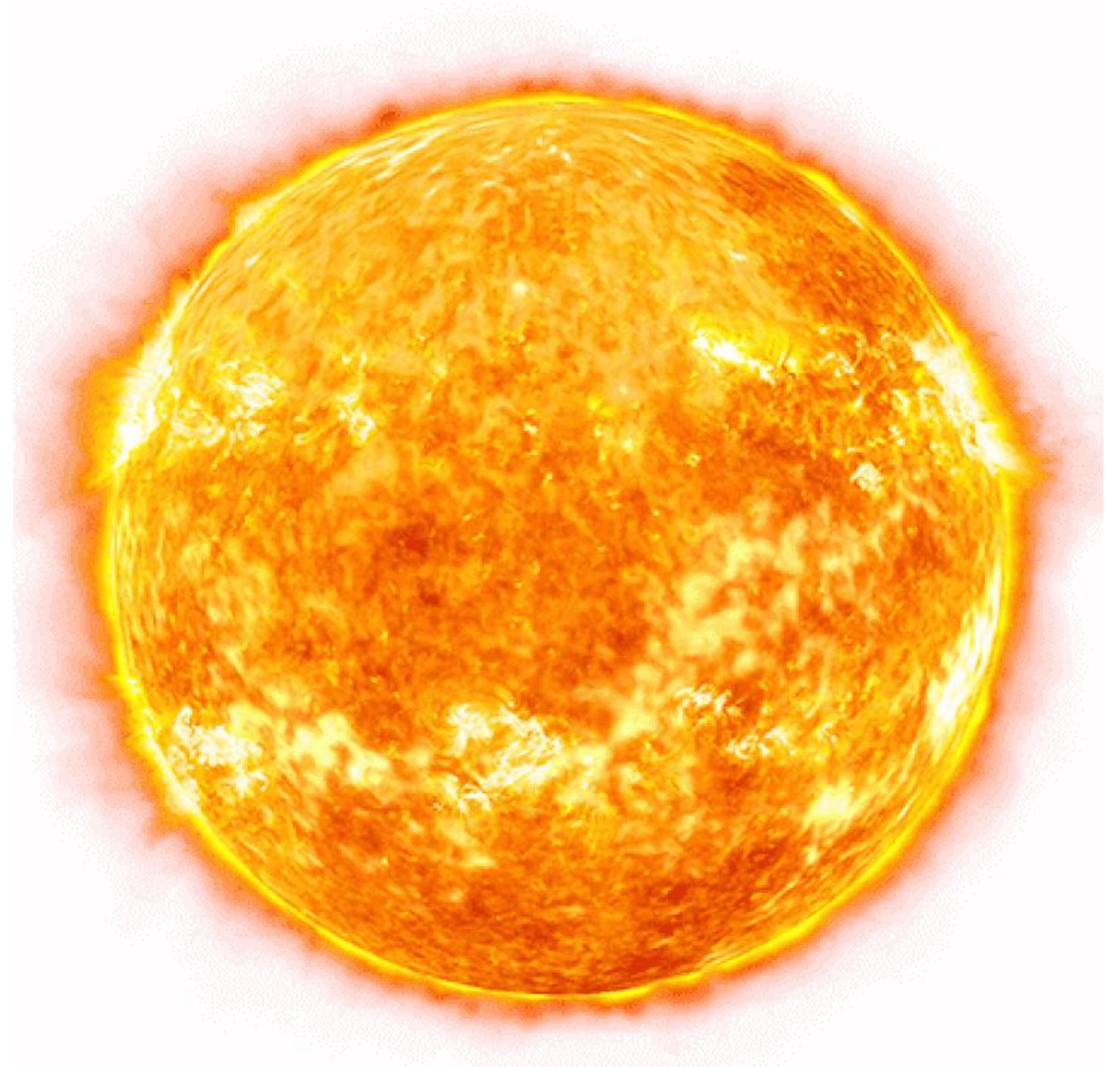
$$\left. \frac{dw}{dt} \right|_{\vec{z}=\vec{z}_i} = -V_i \left(\frac{dE}{dt} \frac{\partial f_0}{\partial E} + \frac{dP_\varphi}{dt} \frac{\partial f_0}{\partial P_\varphi} \right)_{\vec{z}=\vec{z}_i}, w(\vec{z} = \vec{z}_i, t = 0) = 0$$

- In δf method, $w_i = V_i \delta f_i$ with w_i weight of i-th particle. f_0 is given analytically and represents off-axis slowing down NBI distribution

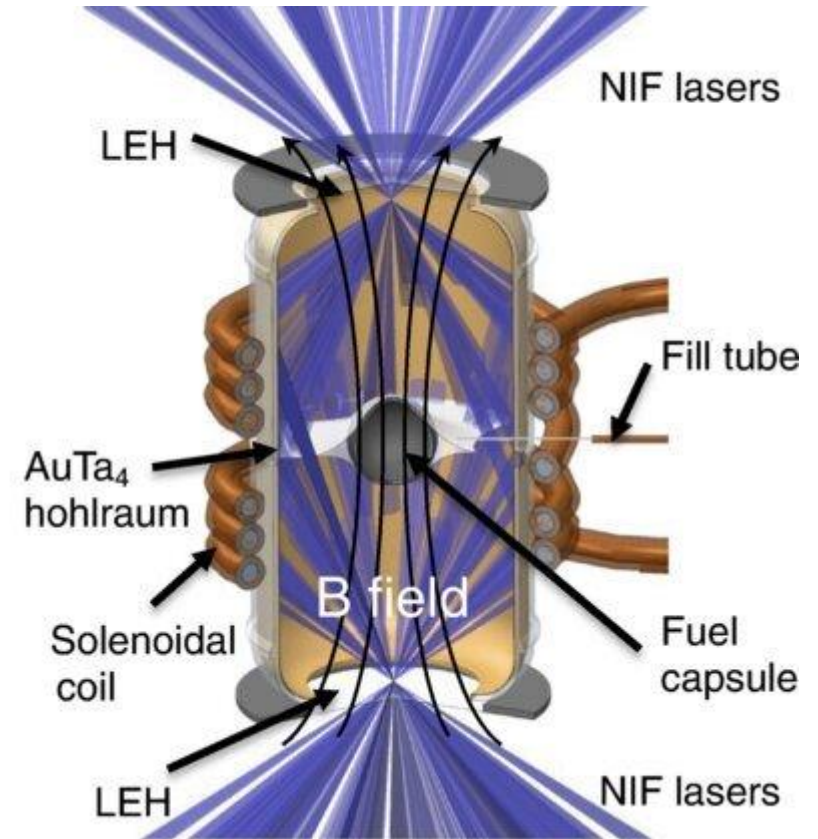
$$p_{EP} = \beta_{EP} \frac{B_0^2}{2\mu_0} \exp \left[- \left(\frac{\Psi_N - \Psi_{N_0}}{\sigma_{\Psi_N}} \right)^2 \right]$$
$$I(v, \Lambda) = \frac{1}{v^3 + v_{\text{crit}}^3} \frac{1}{2} \text{erfc} \left(\frac{v - v_{\text{birth}}}{\Delta_v} \right) \exp \left[- \left(\frac{\Lambda - \Lambda_0}{\Delta_\Lambda} \right)^2 \right]$$

- Gravitational confinement
- Inertial confinement
- Magnetic confinement

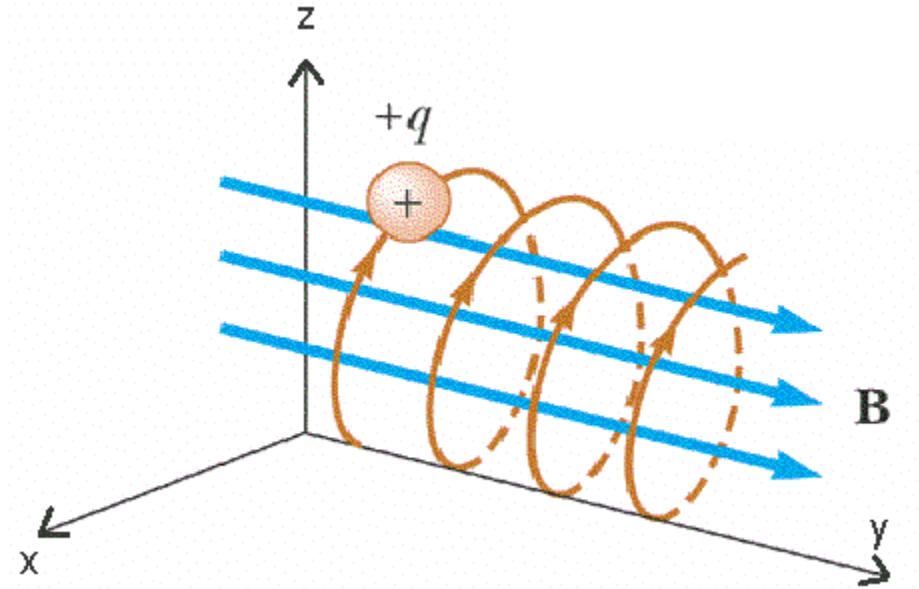
- Gravitational confinement
- Inertial confinement
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- Gravitational confinement
- Inertial confinement
- Magnetic confinement

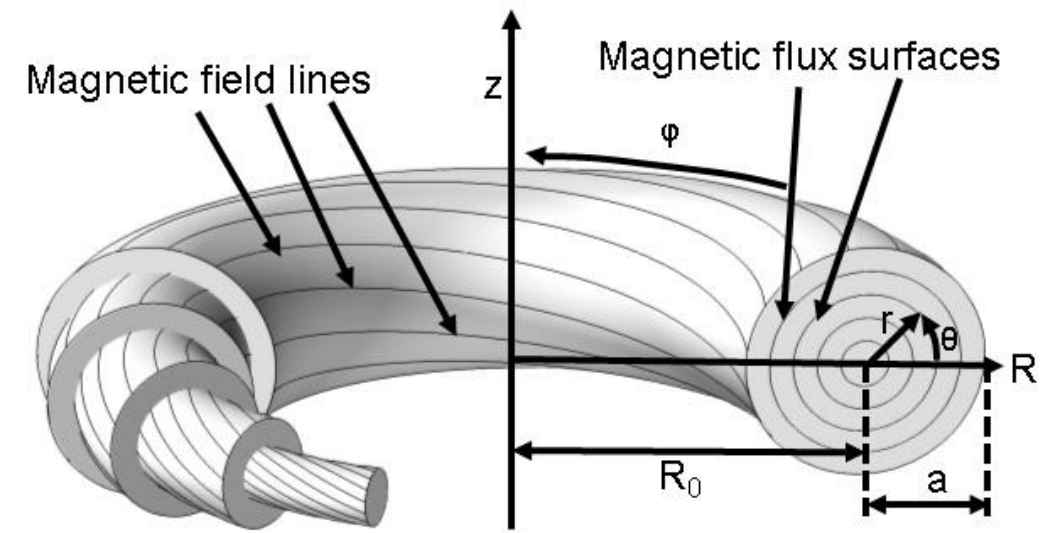


- Gravitational confinement
- Inertial confinement
- Magnetic confinement



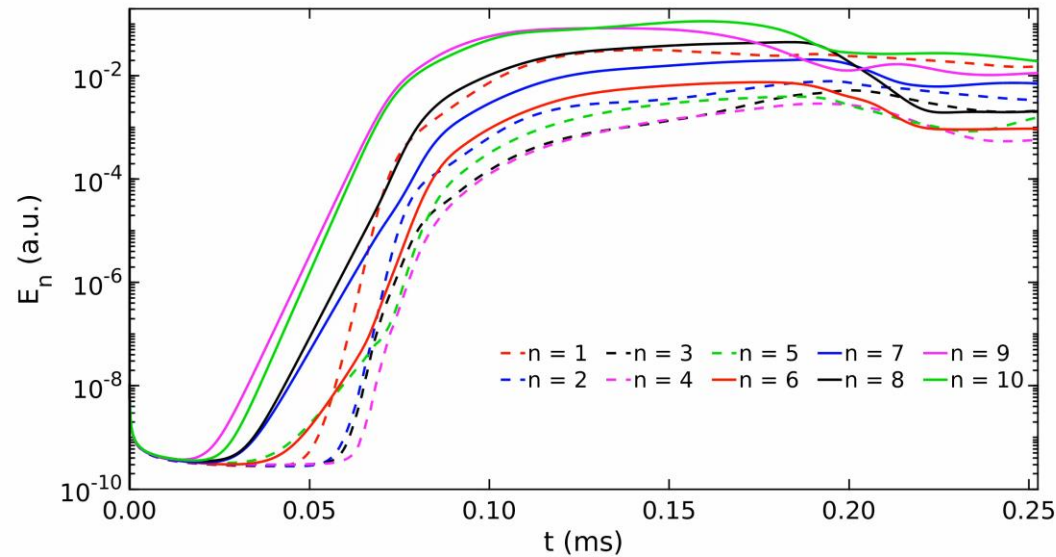
Coordinates in a tokamak

- (R, φ, z) or (r, φ, θ) coordinates
- Poloidal magnetic flux Ψ labels magnetic surfaces
- Normalized poloidal flux $\Psi_N = \frac{\Psi - \Psi_0}{\Psi_a - \Psi_0}$, and $\rho_{\text{pol}} = \sqrt{\Psi_N}$
- Safety factor profile $q = \frac{\text{number of toroidal turns}}{\text{number of poloidal turns}}$

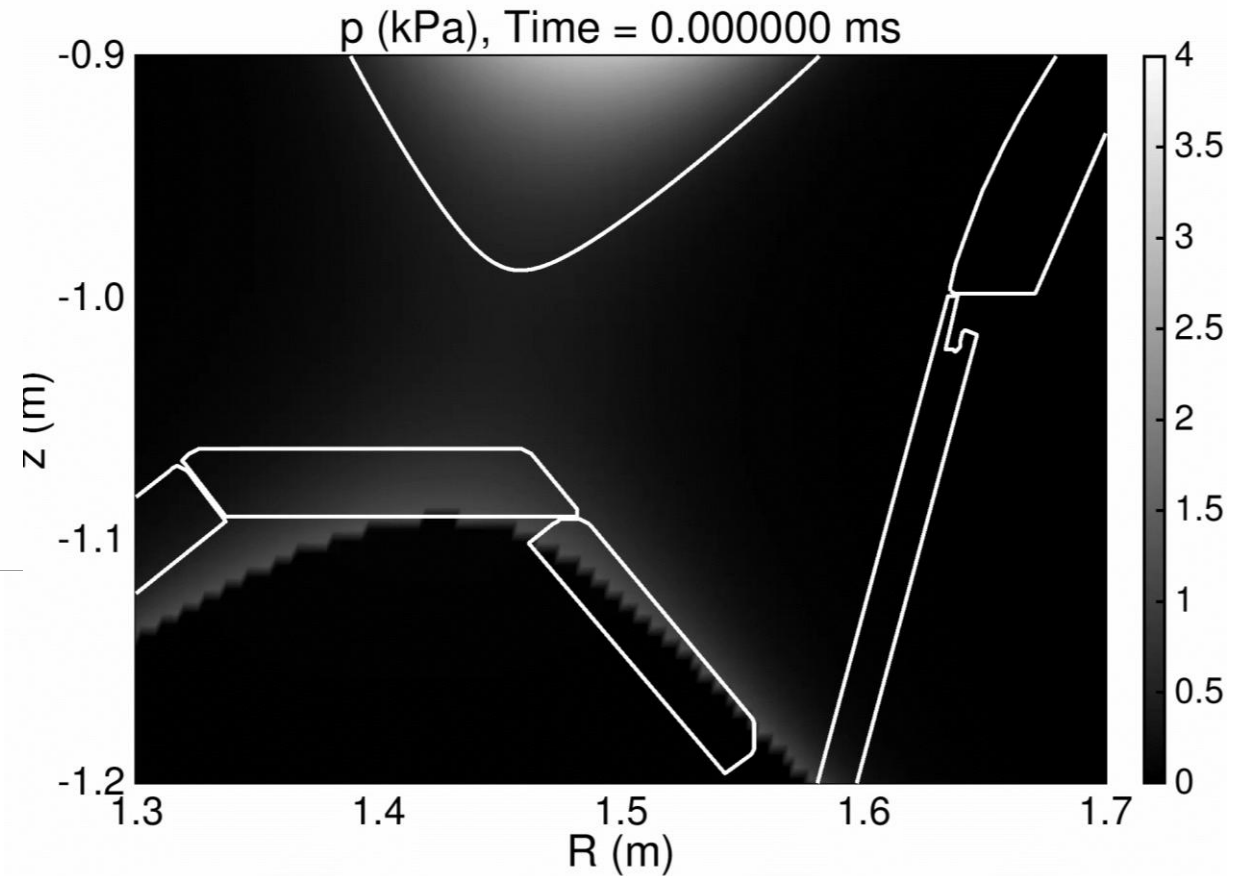


A. Lebschy, PhD Thesis, 2018.

Loss of drive in nonlinear stage (2/2)



- Modes saturate and filaments extend into SOL
- Filaments arrive to outer divertor



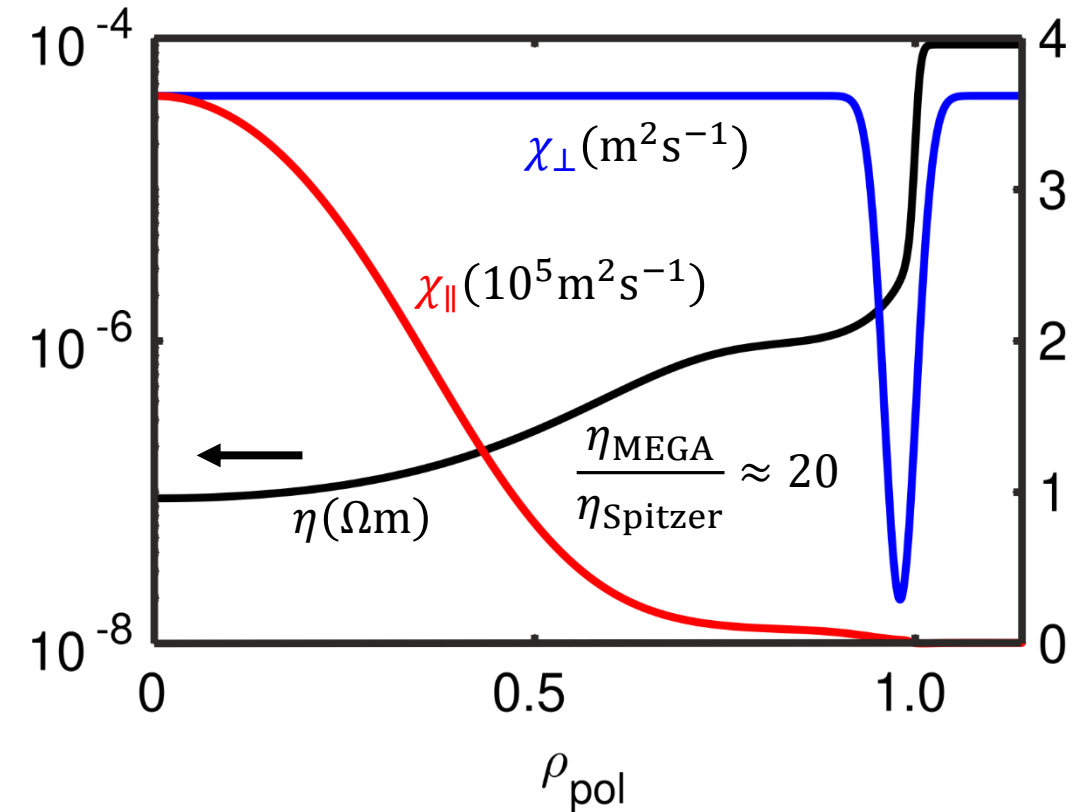
Energy principle: MHD stability analysis of the plasma

- Stability test for arbitrary plasma displacement $\vec{\xi}$
- Potential energy: $\Delta W[\vec{\xi}^*, \vec{\xi}] = -\frac{1}{2} \int d\vec{x} \vec{\xi}^* \cdot \vec{F} \cdot \vec{\xi} \rightarrow$ when $\Delta W > (<) 0$ plasma is stable (unstable)
- Destabilization comes from plasma region

$$\Delta W_F = \frac{1}{2} \int_F d\vec{x} \left[\underbrace{\frac{|B_{1\perp}|^2}{2\mu_0}}_{\text{Magnetic field energy}} + \underbrace{\frac{B_{0\perp}^2}{2\mu_0} |\nabla \cdot \vec{\xi}_\perp + 2\vec{\xi}_\perp \cdot \vec{\kappa}|^2}_{\text{Compression of equilibrium field and plasma}} + \underbrace{\gamma p_0 |\nabla \cdot \vec{\xi}|^2}_{\text{Adiabatic compression}} - \underbrace{2(\vec{\xi}_\perp \cdot \nabla p_0)(\vec{\kappa} \cdot \vec{\xi}^*)}_{\text{Pressure driven}} - \underbrace{\frac{j_{0\parallel}}{B_0} (\vec{\xi}_\perp^* \times \vec{B}_0) \cdot \vec{B}_1}_{\text{Current driven}} \right]$$

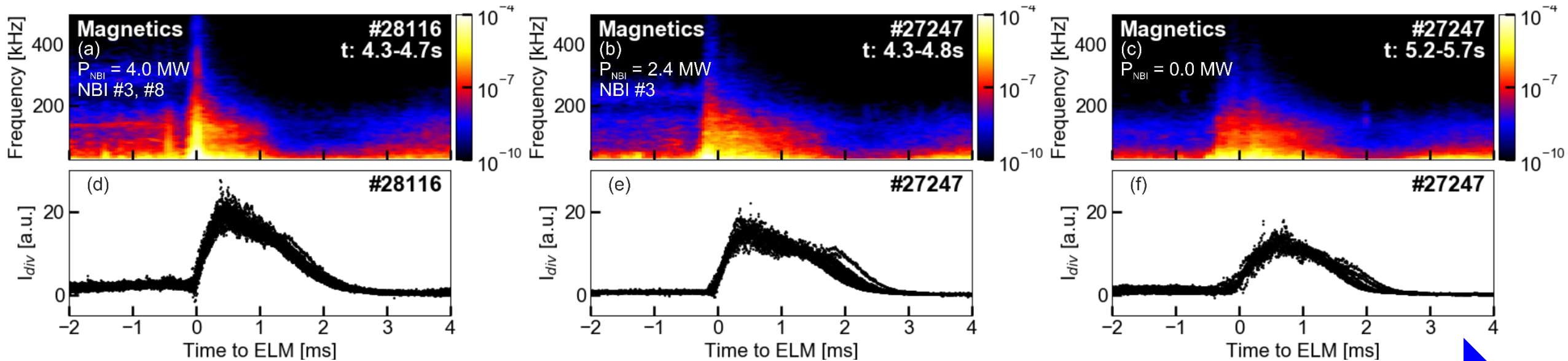
Transport coefficient in the simulations

- Temperature-dependent viscosity and resistivity using **Spitzer's law**
- Ad-hoc profiles for perpendicular thermal diffusivity and particle diffusivity are used to mimic **ETB**^[1]
- Temperature-dependent parallel thermal diffusivity using **Braginskii's law**



[1] E. Viezzer *et al.*, Plasma Phys. Control. Fusion **62**, 024009 (2020)

In experiments, ELM looks remarkably different with and w/o fast-ions



High fast-ion content

Low/zero fast-ion content

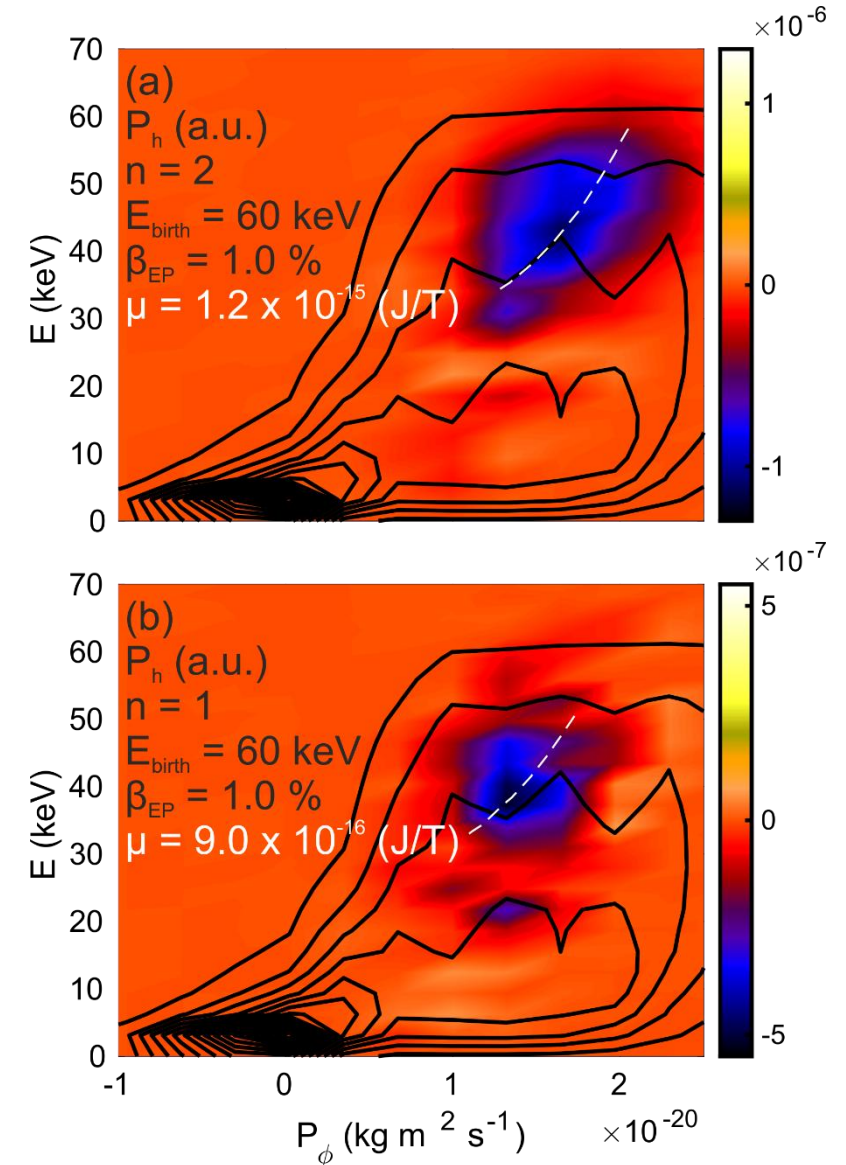
- ELM-synchrhonized spectrogram strongly depends on fast-ion content
 - With larger fast-ion content (low collisionality), ELM crash lasts shorter
 - Magnetic burst at $f > 300$ kHz enhanced with larger fast-ion content

Power transfer with low-n mode numbers

- $P_h \propto \vec{E} \rightarrow$ We can calculate P_h for $\vec{E}_n$
- Phase-space structures aligned with resonant condition^[1]

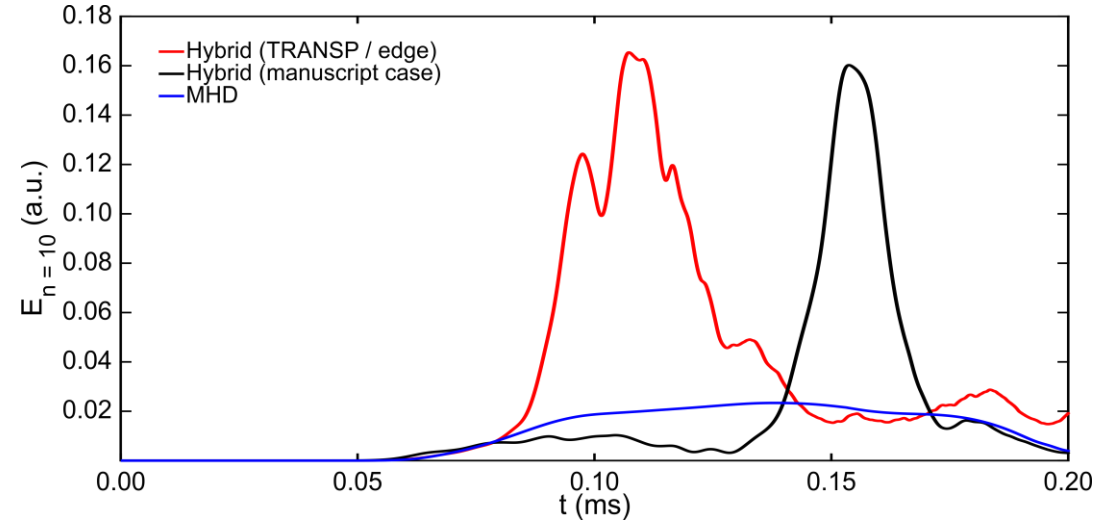
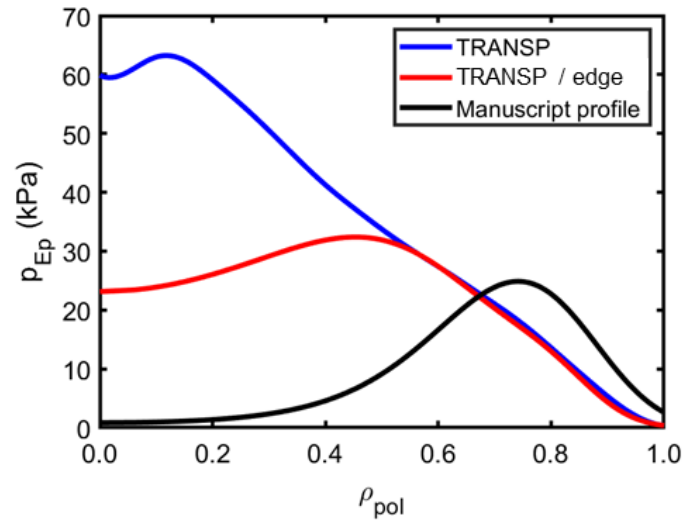
$$\omega_n - n\omega_\phi - p\omega_\theta \approx 0$$

- Resonant energy exchange between low-n modes and fast-ions observed

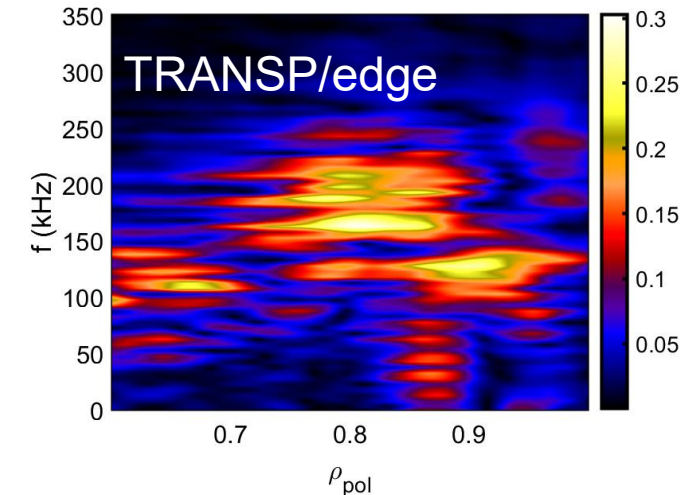


[1] W.W. Heidbrink *et al.*, Phys. Plasmas **27**, 030901 (2020)

Hybrid kinetic-MHD simulations with TRANSP predicted profiles show similar observations

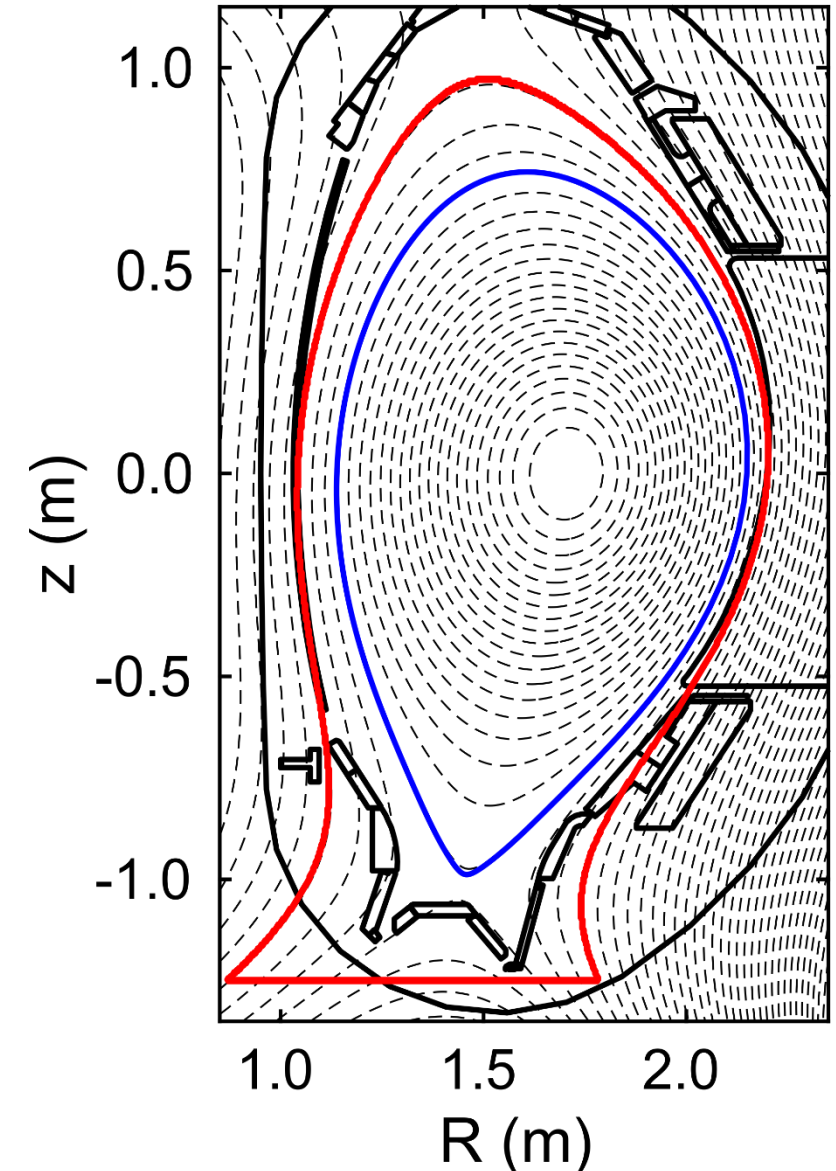


- Fast-ion pressure profile used in MEGA slightly overestimates fast-ion pressure at plasma edge
- Simulations with a TRANSP/edge-like profile still show similar mode energy evolution and impact on the spatial-frequency pattern



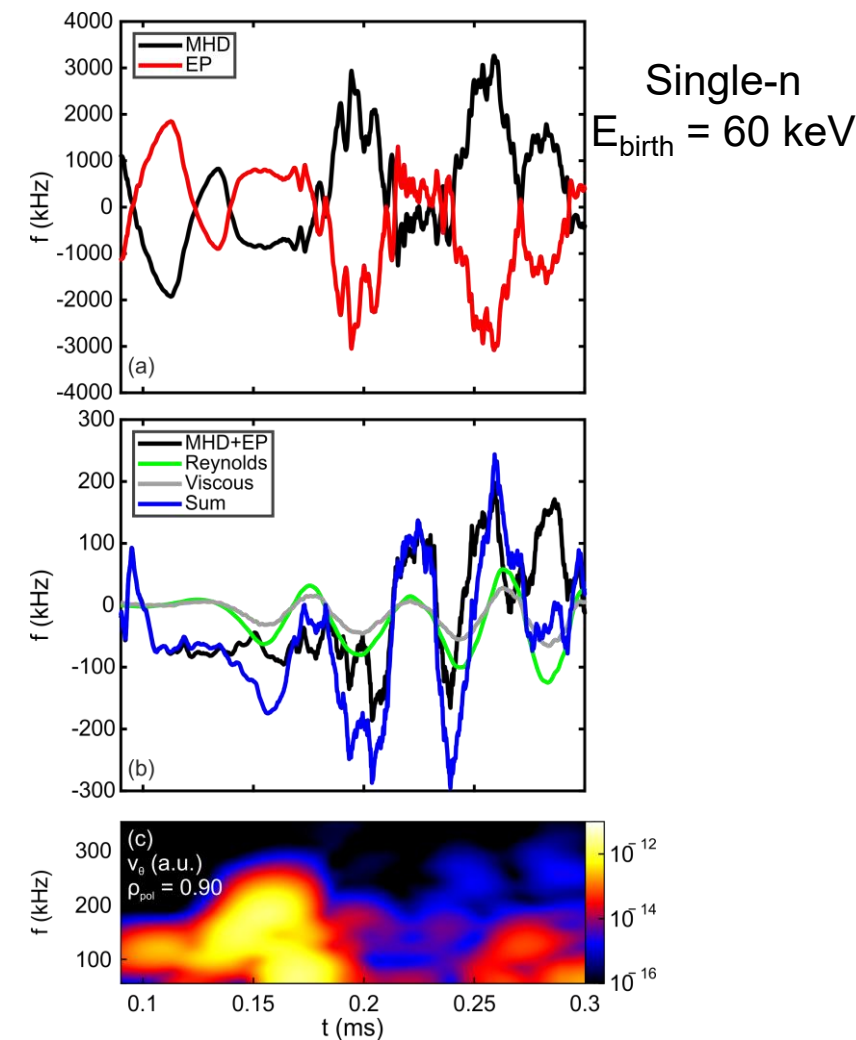
Boundary conditions in the simulations

- $\vec{u} = \vec{0}, \delta\rho = \delta p = 0$ outside area enclosed by red line (binary mask) seen in right figure
- At boundary of simulation domain, $\delta\vec{B} \cdot \hat{n} = 0$ and $\vec{E} \times \hat{n} = \vec{0} \rightarrow$ perfect conducting wall boundary condition
- Distribution function of particles $f = 0$ at boundary of simulation domain



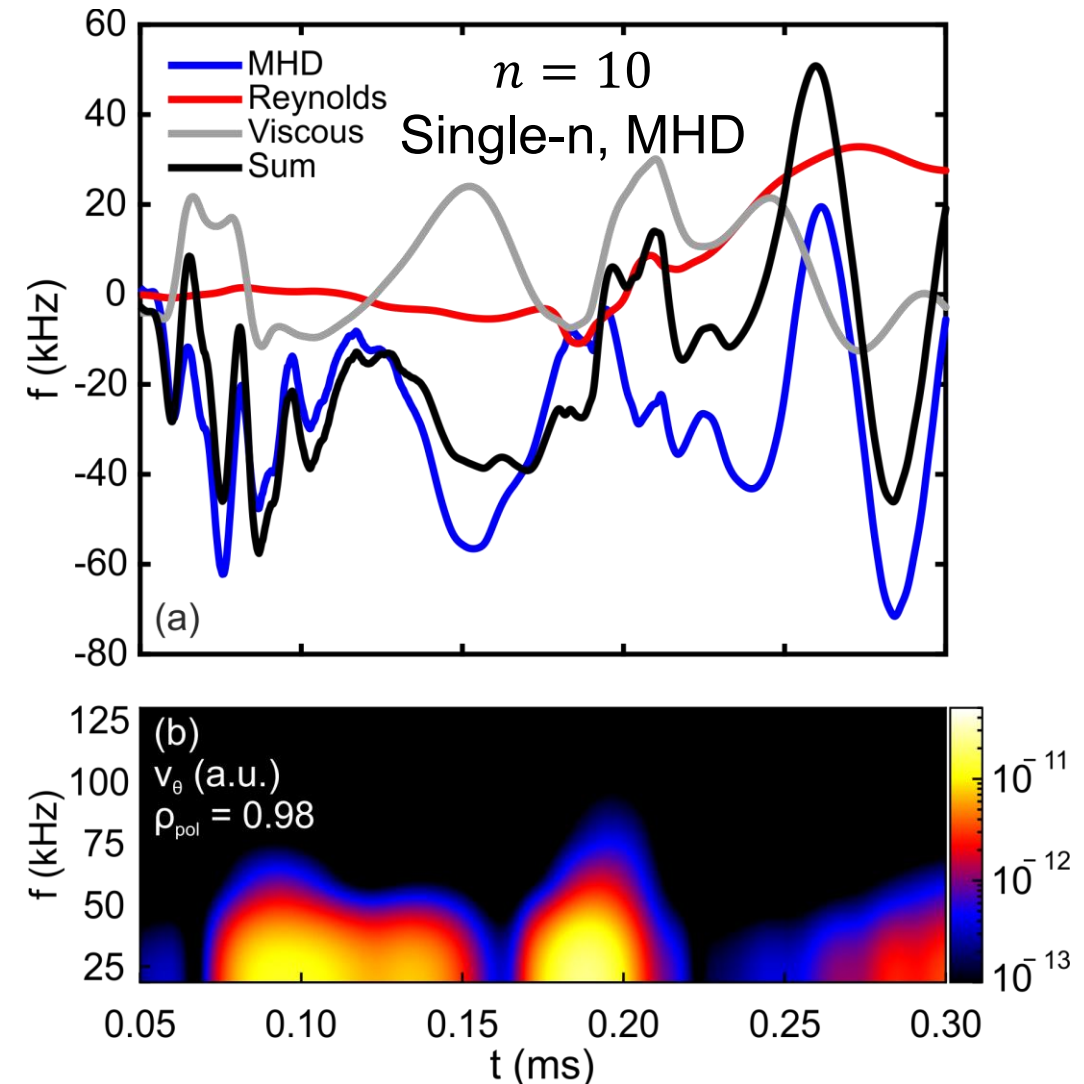
Imbalance of forces in hybrid kinetic-MHD simulations of ELMs

- In hybrid kinetic-MHD simulations, the frequency is significantly boosted
- Main forces driving mode frequency are MHD and energetic particle forces
- Reynolds force plays a key role in the total contribution of forces to mode frequency



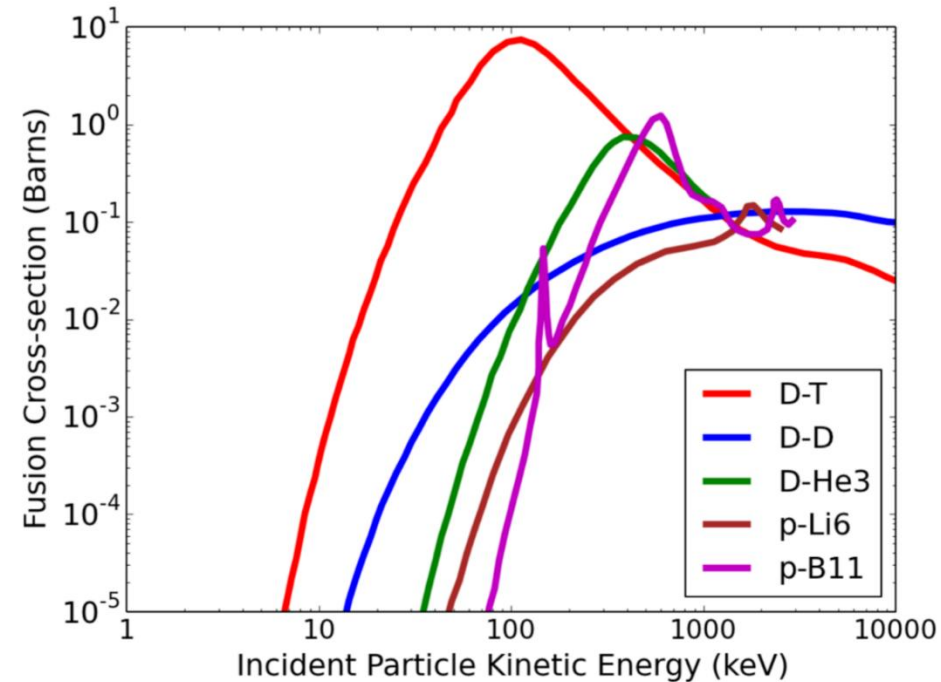
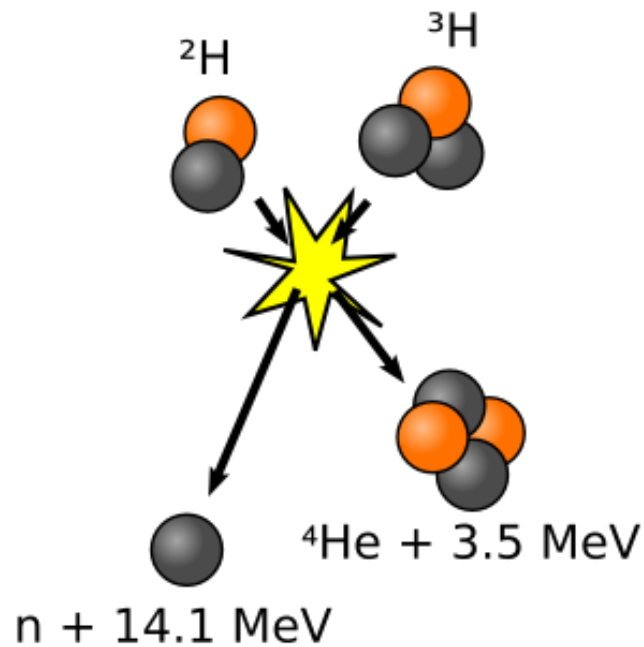
Mode rotation brought by imbalance of forces in the MHD momentum equation

- **MHD force** ($= -\nabla p + \vec{j} \times \vec{B}$) determines mode frequency
- **Viscous force** [$= -\nabla \times (v\rho\nabla \times \vec{u}) + \frac{4}{3}\nabla(v\rho\nabla \cdot \vec{u})$] generally damps mode frequency in nonlinear phase
- **Reynolds force** [$= -\nabla \cdot (\rho\vec{u}\vec{u})$] play major role long after ELM in mode frequency



Nuclear fusion reaction on Earth

- Mass deficit leads to large energy release due to $E = mc^2$
- D-T reaction has the largest cross section for achievable temperatures



The ignition condition

- We need plasma conditions such that $Q_{\text{out}} \gg Q_{\text{in}}$

The ignition condition

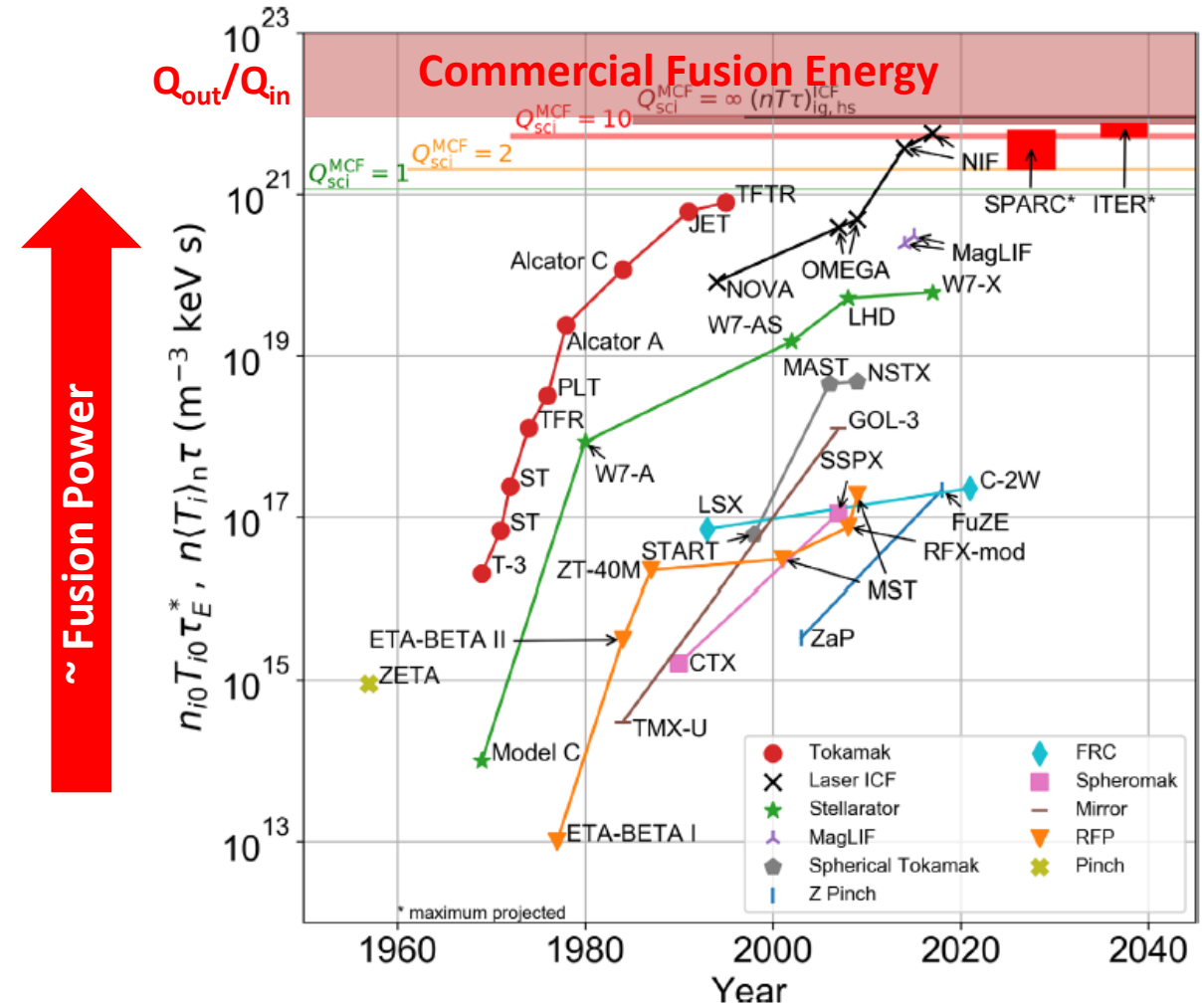
- We need plasma conditions such that $Q_{\text{out}} \gg Q_{\text{in}}$
- Lawson criteria^[1]

$$nT\tau_E > 5 \times 10^{21} \text{ m}^3 \text{ s keV}$$

n : plasma density

T : plasma temperature

τ_E : energy confinement time



Wurzel and Hsu, Phys. Plasmas **29** 062103 (2022)

[1] J.D. Lawson *et al.*, Proc. Phys. Soc. B **70**, 6 (1957)

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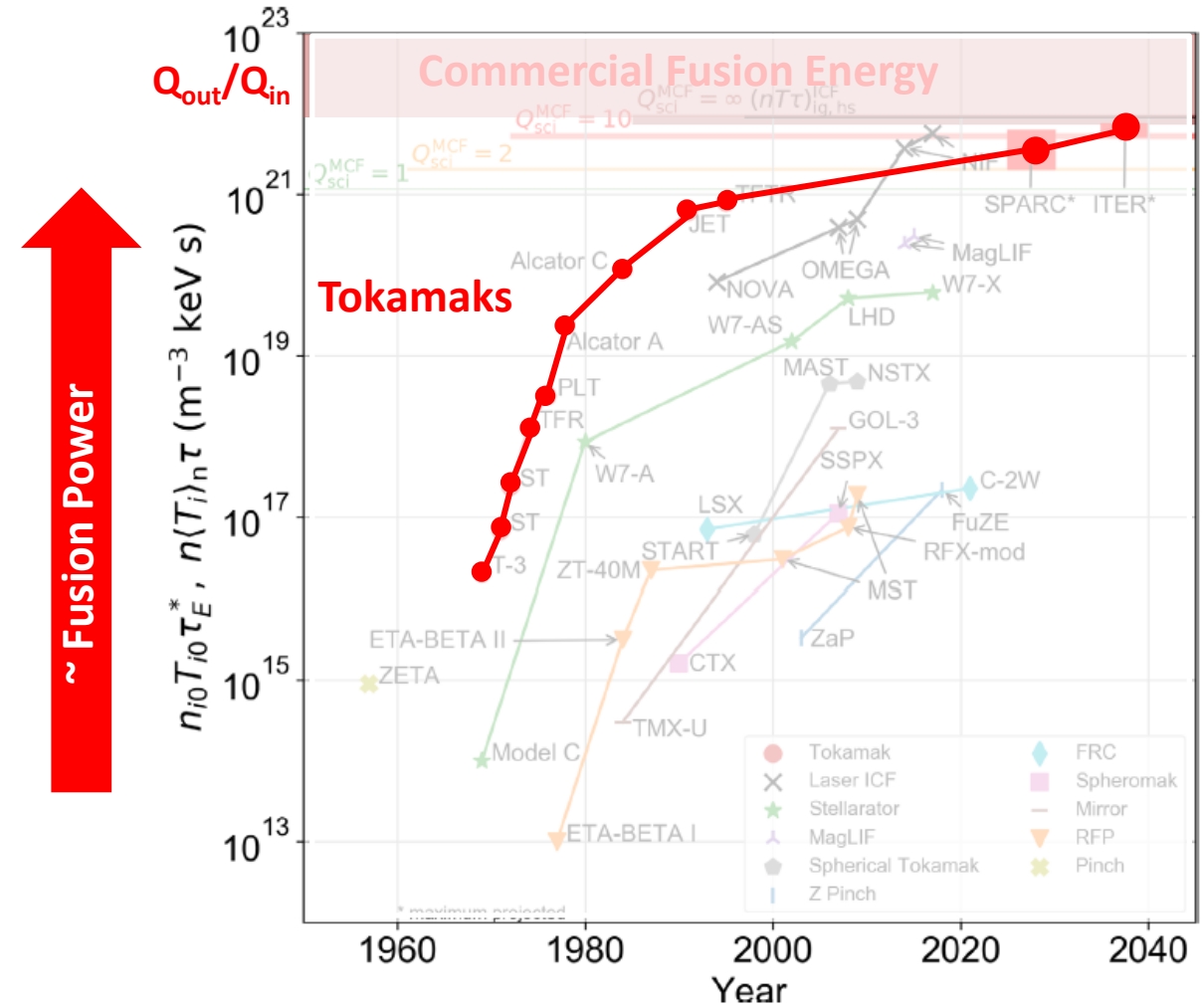
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- Magnetic confinement (tokamaks) are the most advanced concept



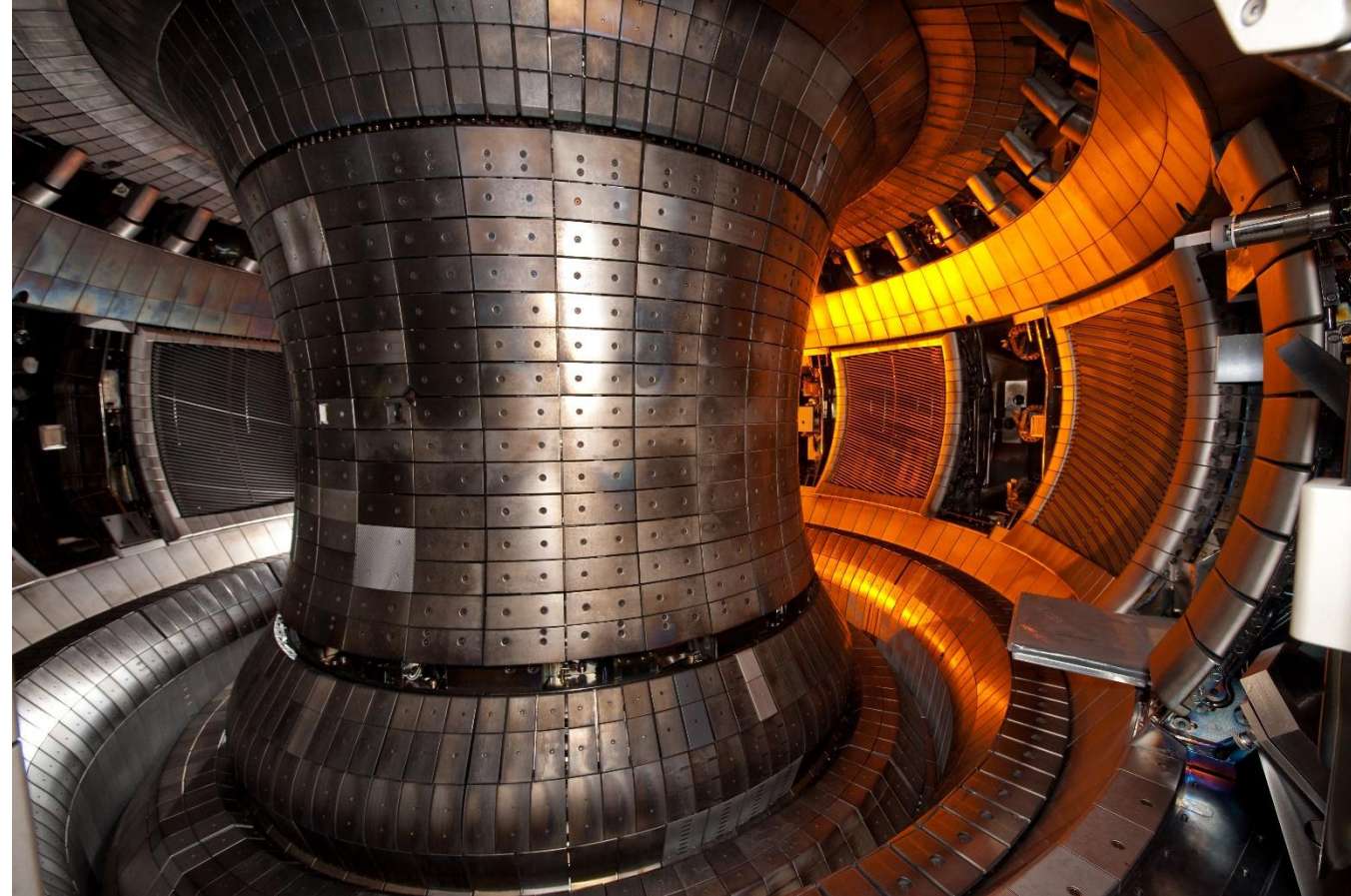
Wurzel and Hsu, Phys. Plasmas **29** 062103 (2022)

[1] J.D. Lawson *et al.*, Proc. Phys. Soc. B **70**, 6 (1957)

ASDEX Upgrade tokamak (AUG)

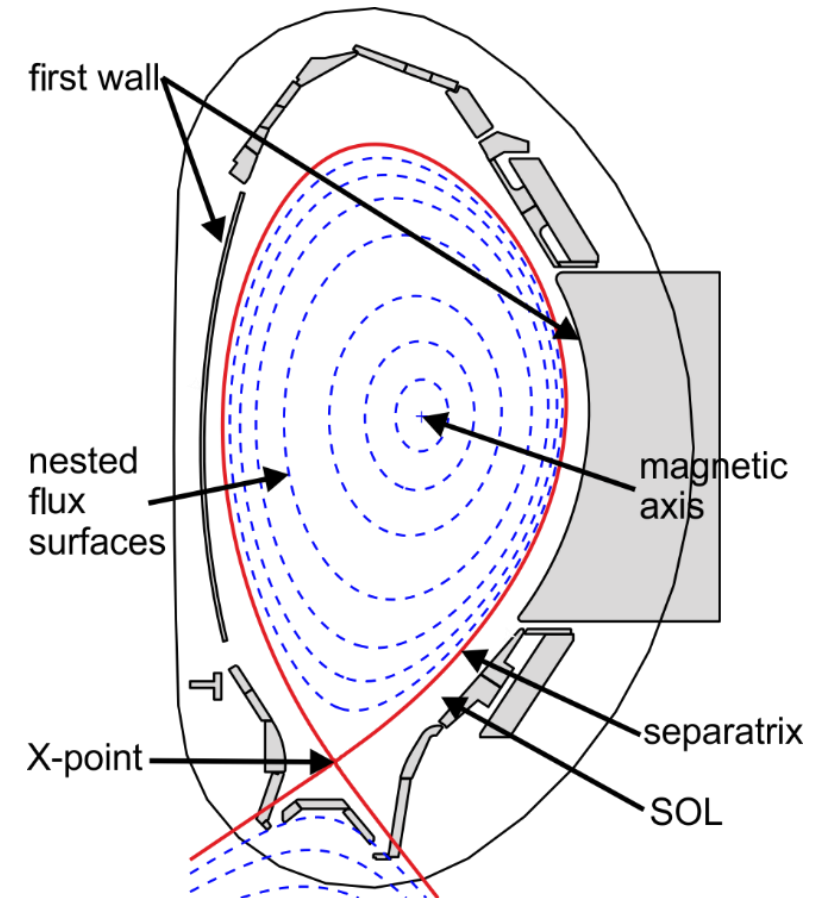
- Medium-sized tokamak located at Max-Planck Institute for Plasma Physics (Garching, Germany)
- Main parameters of AUG

Major radius R_0	1.65 m
Minor radius a	0.5 m
Plasma current I_p	0.4 – 1.2 MA
Toroidal magnetic field B_t	1.8 – 2.8 T
External heating power	up to 30 MW
Main ion	D, H, He



The plasma shape in a tokamak

- Separatrix: boundary of confined plasma
- Nested flux surfaces are labeled by poloidal magnetic flux Ψ
- Normalized poloidal flux $\Psi_N = \frac{\Psi - \Psi_0}{\Psi_s - \Psi_0}$, and
 $\rho_{\text{pol}} = \sqrt{\Psi_N}$
- SOL: unconfined region



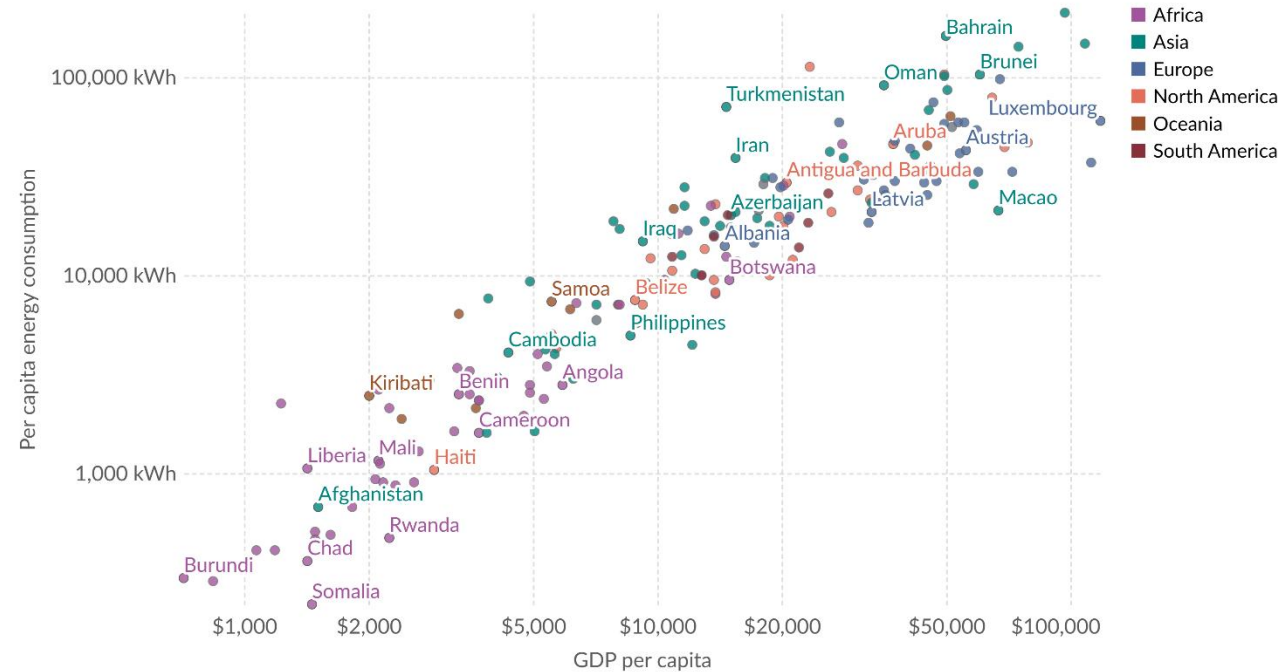
*D.J. Cruz-Zabala PhD thesis

Energy consumption is correlated with wealthiness

Energy use per person vs. GDP per capita, 2022

Our World
in Data

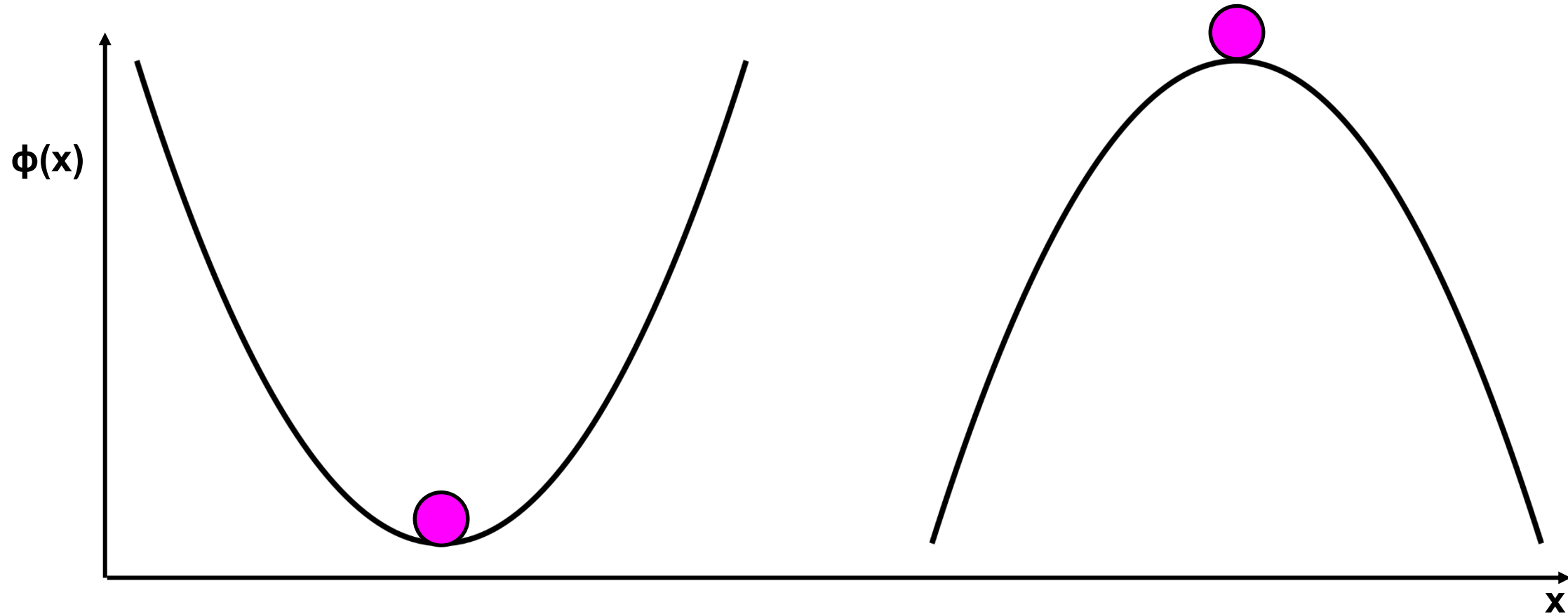
Energy refers to primary energy¹, measured in kilowatt-hours² per person, using the substitution method³. Gross domestic product (GDP) is adjusted for inflation and differences in the cost of living between countries.



Data source: U.S. Energy Information Administration (2023) and other sources

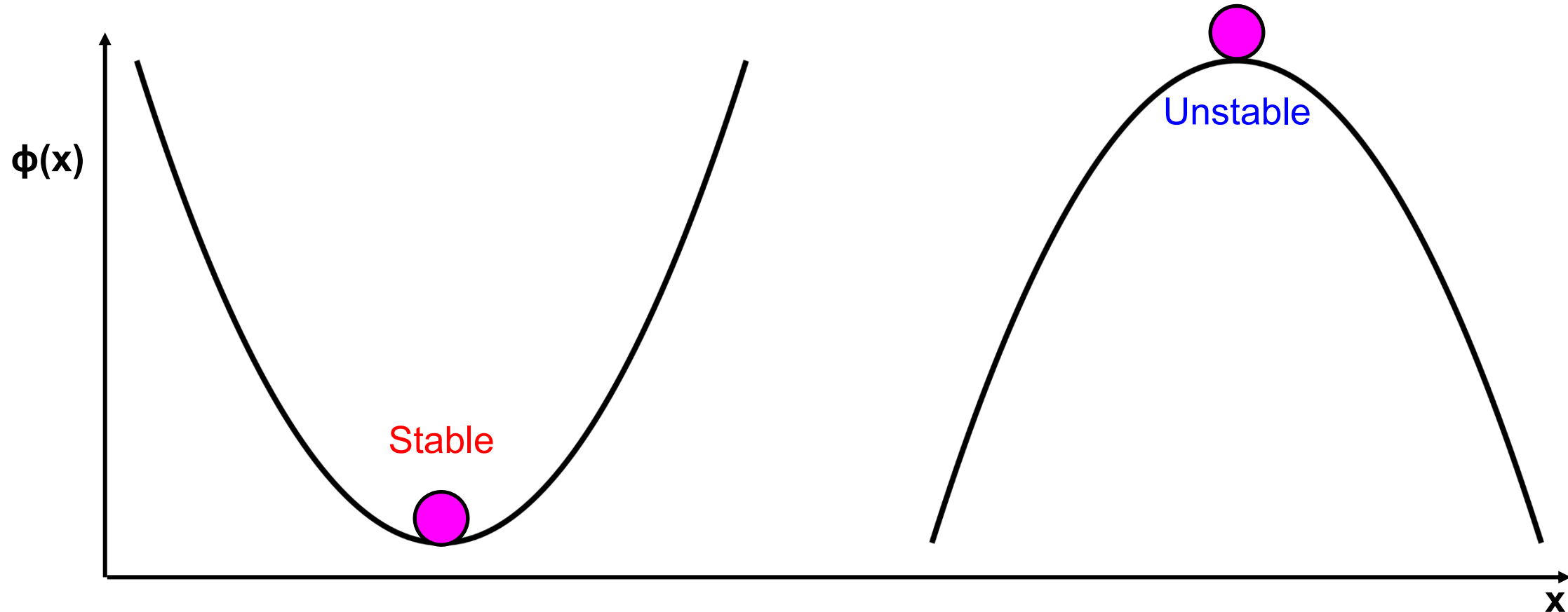
OurWorldinData.org/energy | CC BY

The stability analysis of a particle in classical mechanics



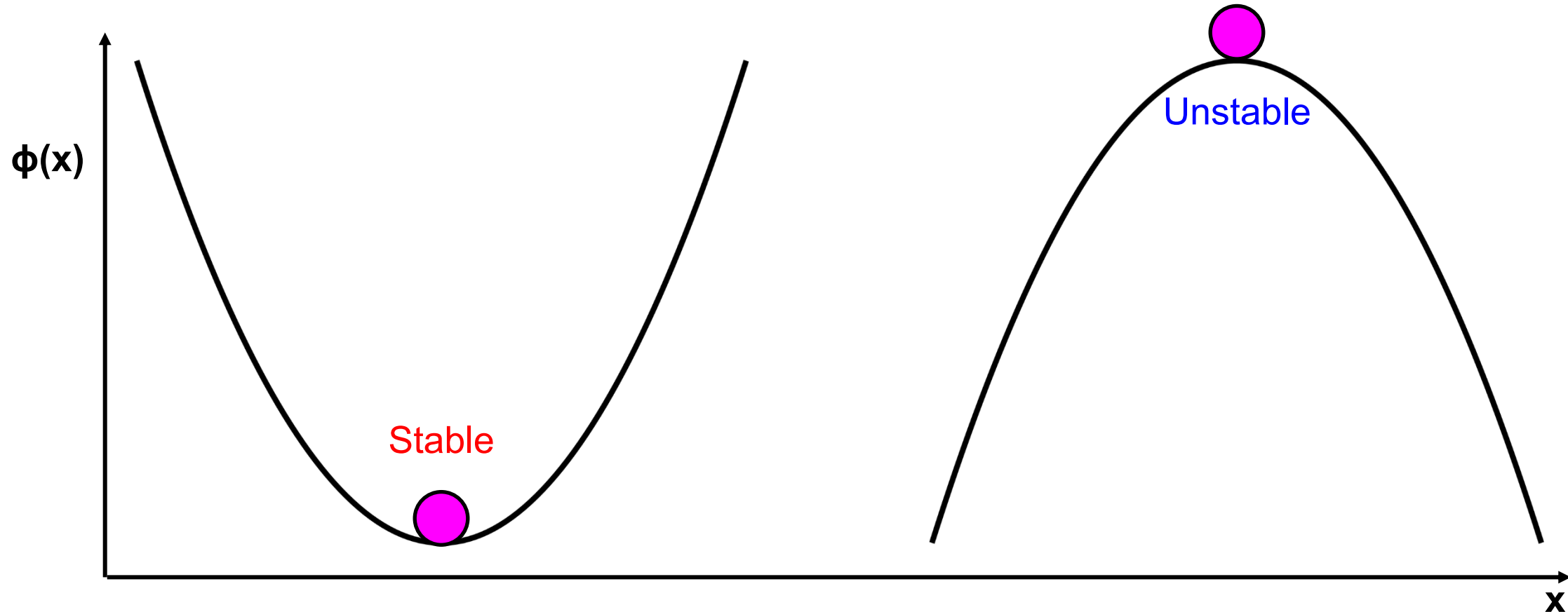
- In classical mechanics, particle stability depends on how the potential $\phi(x)$ is modified by a perturbation around equilibrium

The stability analysis of a particle in classical mechanics



- In classical mechanics, particle stability depends on how the potential $\phi(x)$ is modified by a perturbation around equilibrium

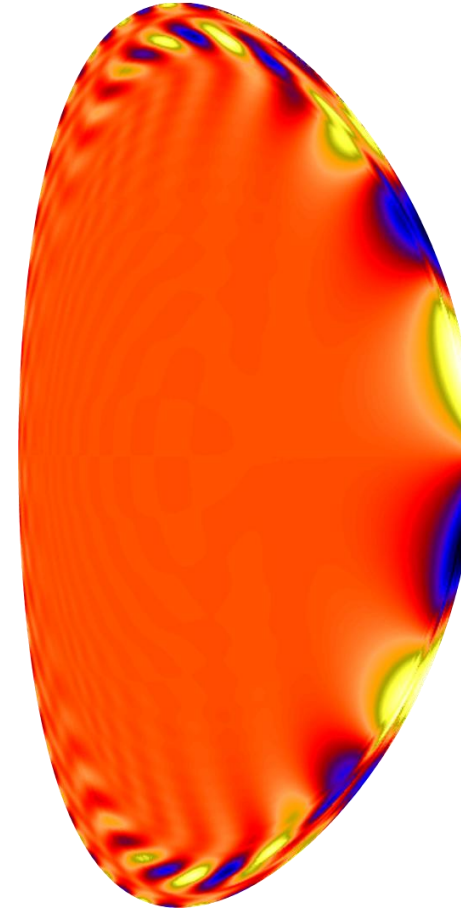
The stability analysis of a particle in classical mechanics



- In classical mechanics, particle stability depends on how the potential $\phi(x)$ is modified by a perturbation around equilibrium
- A similar analysis can be developed in MHD theory

Destabilizing mechanism of MHD instabilities

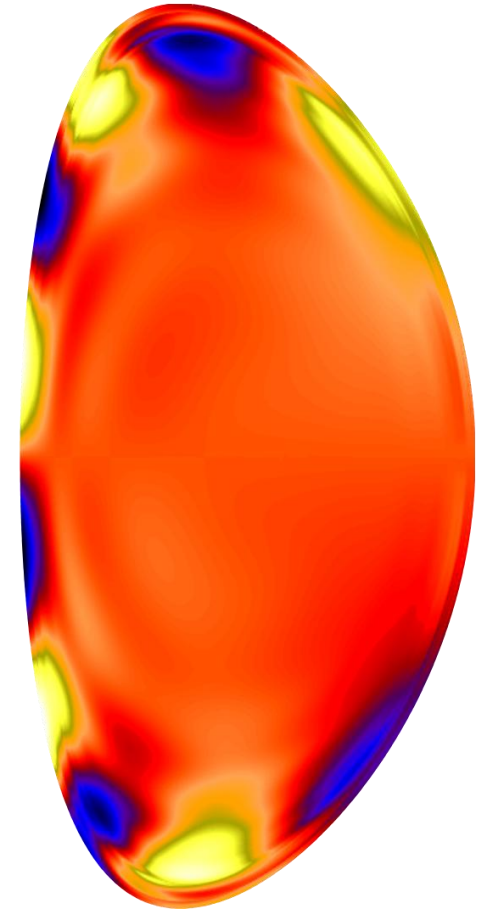
- Pressure-driven instabilities
 - Located in bad curvature region
 - E.g.: ballooning modes



Ballooning

Destabilizing mechanism of MHD instabilities

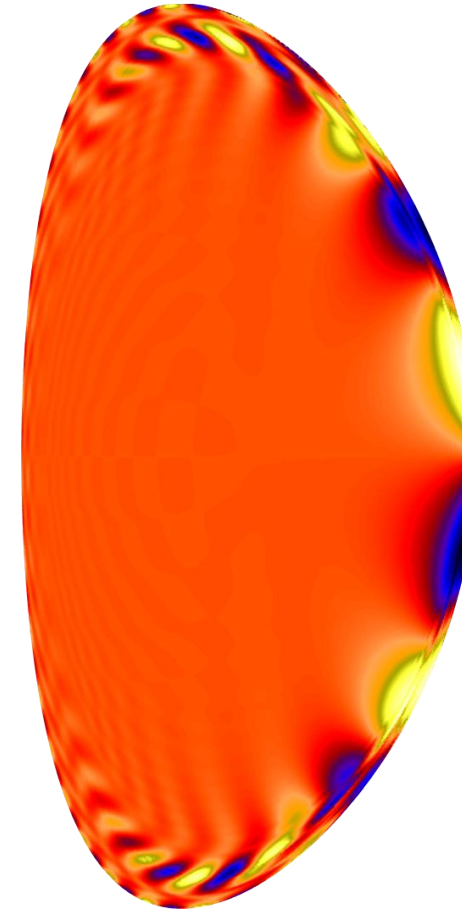
- Current-driven instabilities
 - E.g.: peeling-modes



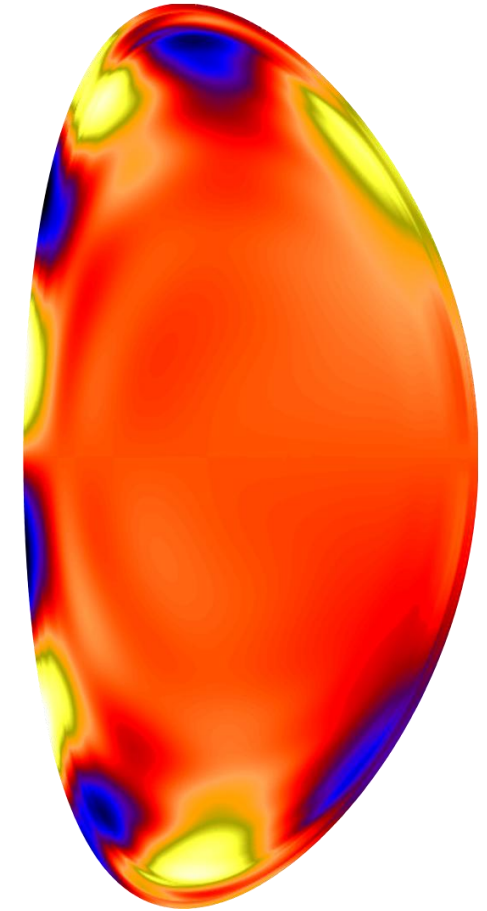
Peeling

Destabilizing mechanism of MHD instabilities

- Pressure-driven instabilities
 - Located in bad curvature region
 - E.g.: ballooning modes
- Current-driven instabilities
 - E.g.: peeling-modes
- ELM onset is described as coupled peeling-balloonning modes^[1]



Ballooning



Peeling

[1] J.W. Connor *et al.*, Phys. Plasmas **5**, 2687 (1998)

- Nonlinear hybrid kinetic-MHD code that calculates the self-consistent interaction between energetic particles and thermal plasma^[1]

[1] Y. Todo *et al.*, Phys. Plasmas **5**, 1321 (1998)

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- **Current coupling** scheme in **MHD momentum equation**

$$\frac{\partial(\rho\vec{u})}{\partial t} = -\nabla \cdot (\rho\vec{u}\vec{u}) - \nabla p + (\vec{j} - \quad) \times \vec{B} - \nabla \times (v\rho\nabla \times \vec{u}) + \frac{4}{3}\nabla \times (v\rho\nabla \times \vec{u})$$

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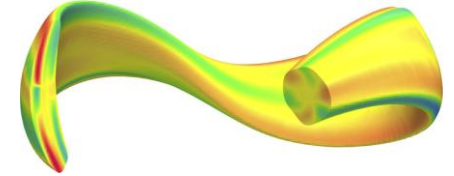
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- Widely used in fusion community

CFQS



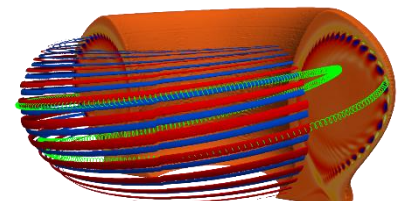
H. Wang et al., NF (2022)

Heliotron-J



P. Adulsiriswad et al., NF (2020)

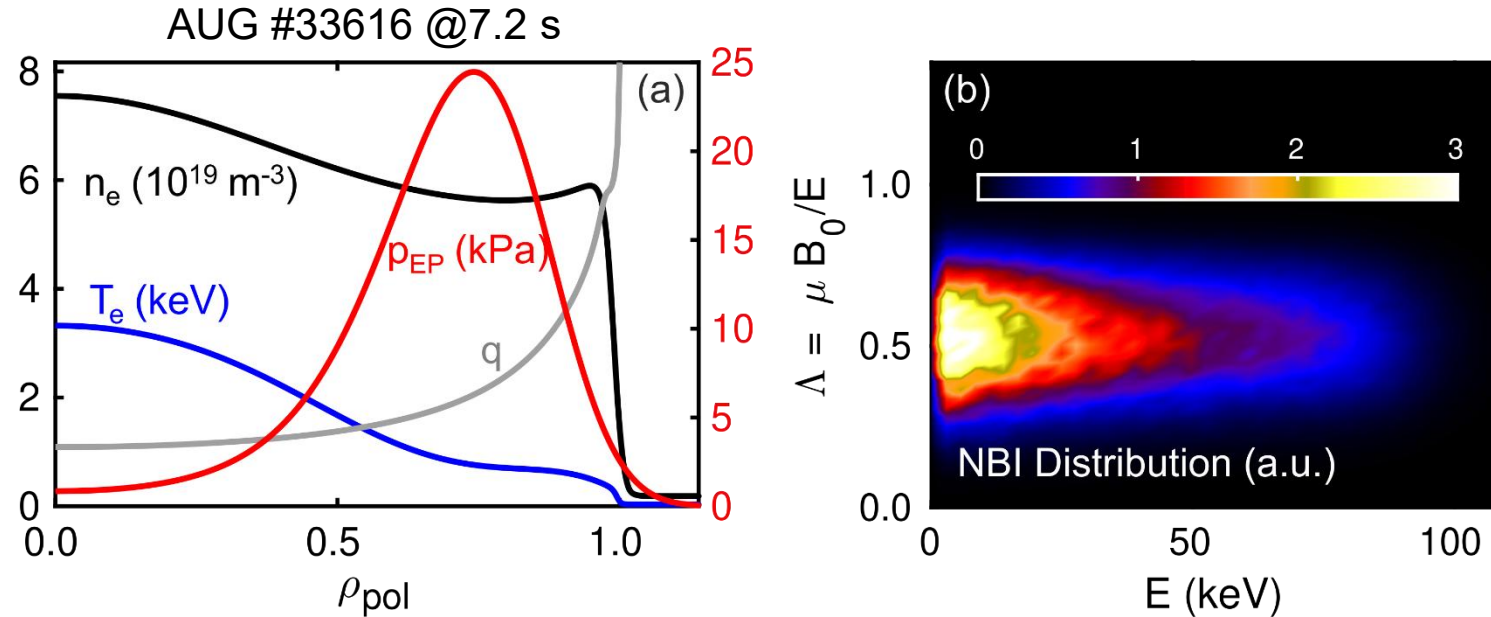
AUG



J. Dominguez-Palacios et al.,
Nat. Phys. (Submitted)

[1] Y. Todo *et al.*, Phys. Plasmas **5**, 1321 (1998)

Realistic initial conditions for bulk plasma and energetic particles



- Thermal plasma: measured kinetic profiles used as initial conditions^[1] and standard MHD model^[2]
- Fast-ions: on/off-axis slowing down NBI distribution

[1] A.F. Mink *et al.*, Nucl. Fusion **58**, 026011 (2018)

[2] Y. Todo *et al.*, Phys. Plasmas **24**, 081203 (2017)

- Time integration: 4th order explicit Runge-Kutta method
- Fluid fields
 - Spatial derivatives: 4th order finite difference method

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Numerical methods of MEGA

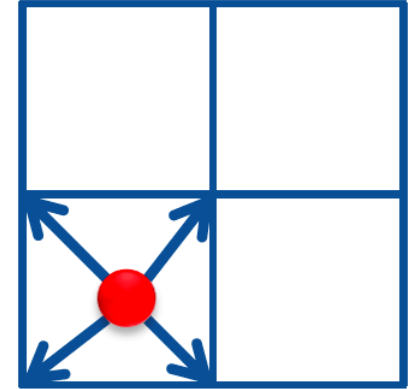
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Information of particle distributions is assigned to grids

Numerical methods of MEGA

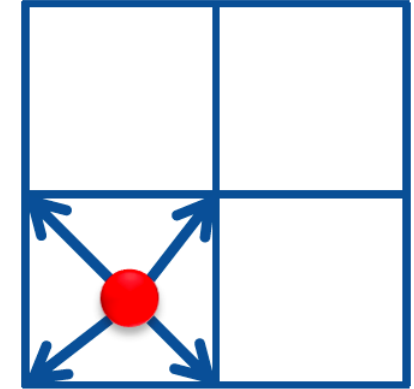
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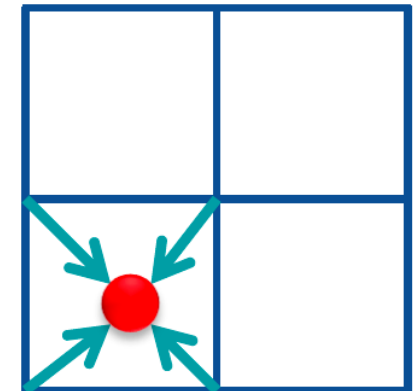
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Forces on a particle are derived from the fields

Numerical methods of MEGA

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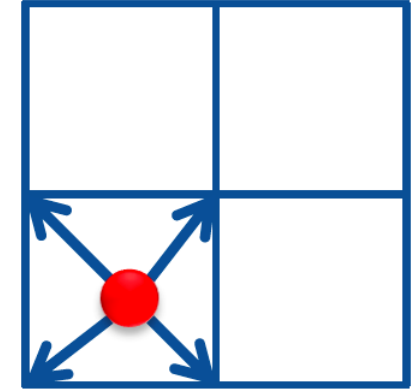
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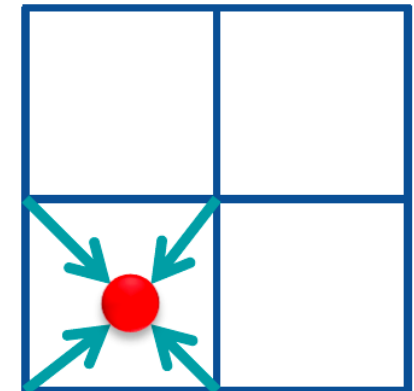
- Particles

→ 1st order particle-in-cell (PIC) with δf or full-f method

→ FLR effects on particle orbits



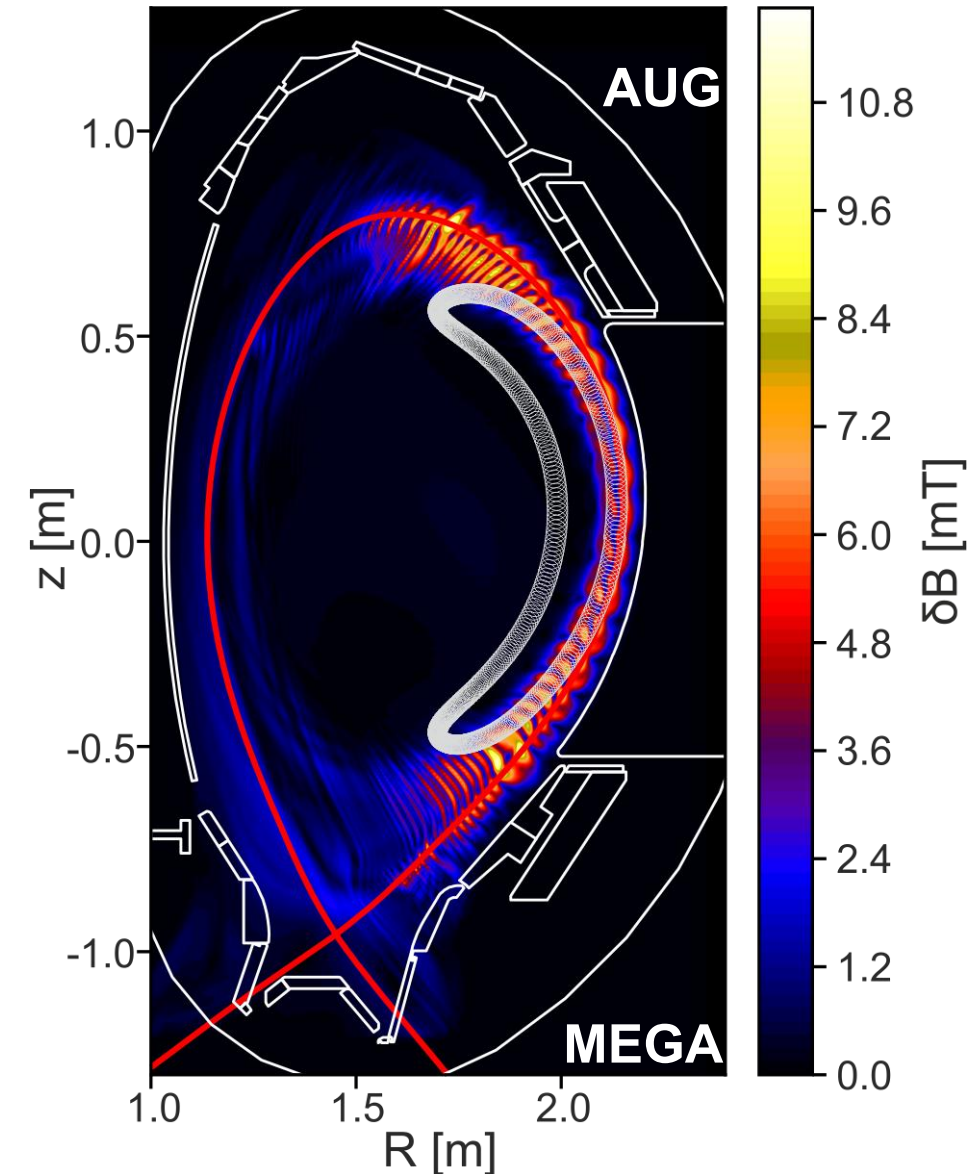
Information of particle distributions is assigned to grids



Forces on a particle are derived from the fields

Simulation domain in MEGA

- SOL and Private Flux Region below X-Point included in simulation domain
 - ELM relevant area
 - Important to study the interaction with fast-ions
- 3D cartesian grids of cylindrical coordinates → inclusion of SOL and PFR depends only on magnetic field topology^[1, 2]

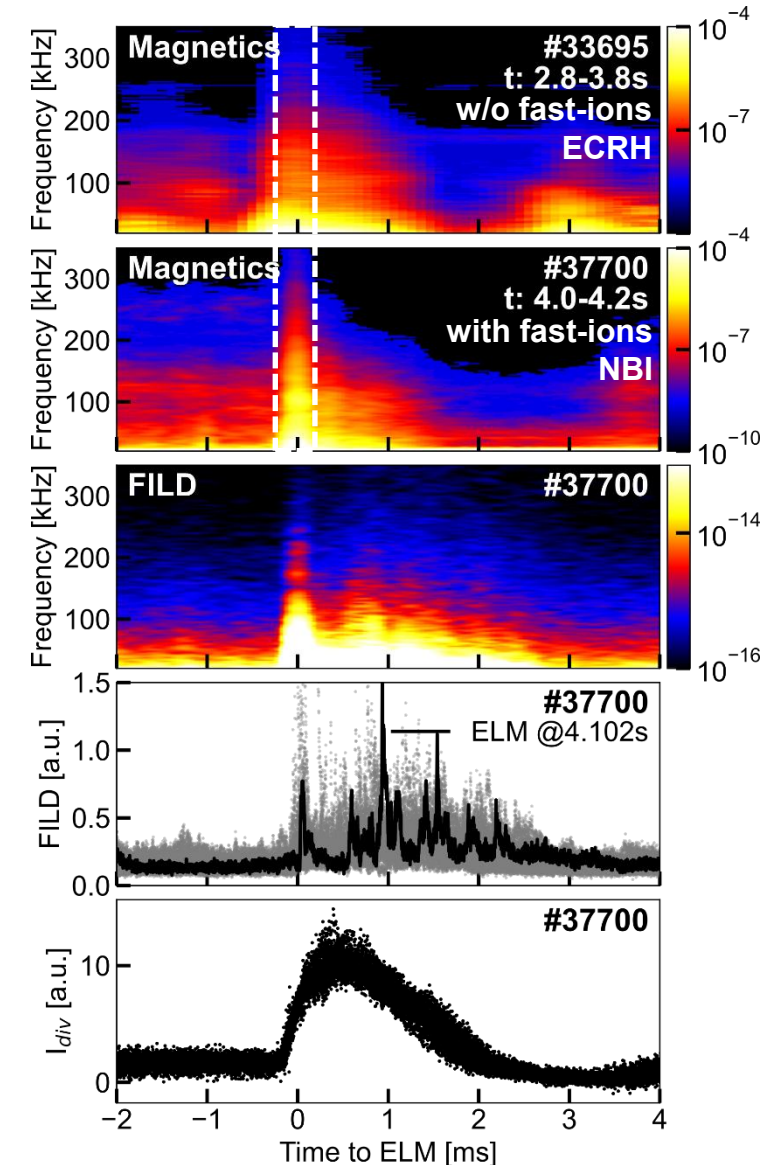


[1] R. Khan *et al.*, Phys. Plasmas **14**, 062302 (2007)

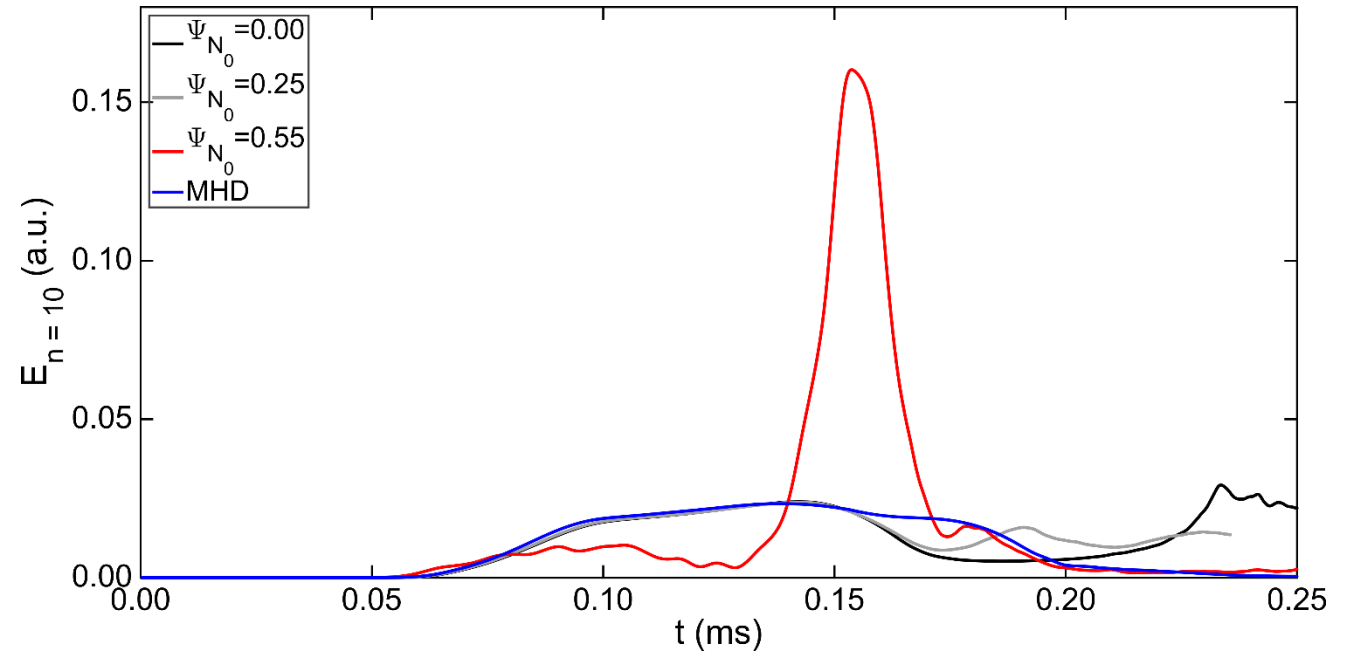
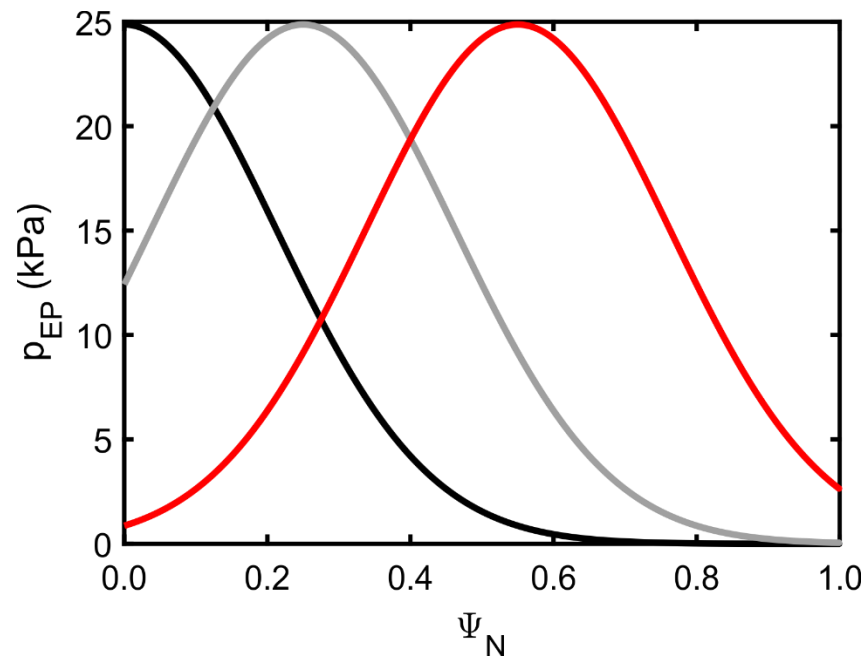
[2] N. Mizuguchi *et al.*, Nucl. Fusion **47**, 579 (2007)

In experiments, ELM looks remarkably different with and w/o fast-ions

- ELM synchronised magnetic spectrograms seem to depend strongly on fast-ion population
 - **Large fast-ion content** in NBI heated discharges with low collisionality lead to abrupt crashes ($<100\ \mu\text{s}$) that are extended to high frequencies ($\sim 300\ \text{kHz}$)
 - Amplitude of low-frequency oscillations ($<200\ \text{kHz}$) decay in $\sim 2\ \text{ms}$
 - In discharges with **no fast-ion** sources and high collisionality, ELM crashes seem to last longer and have lower amplitudes especially at high frequencies
- In NBI heated discharges, FILD and magnetic spectrograms show striking similarities



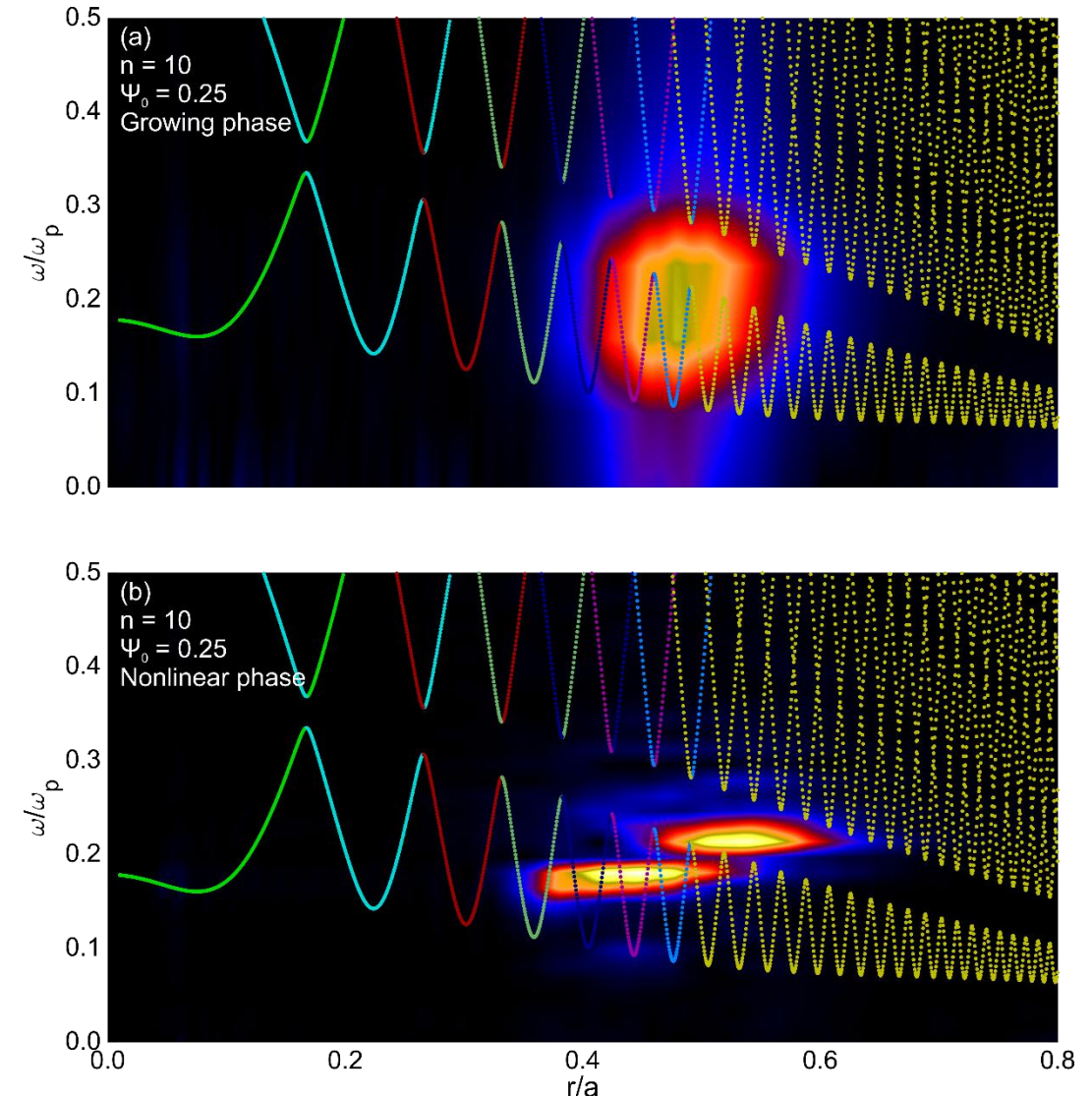
Impact of off-axis peak location on ELM



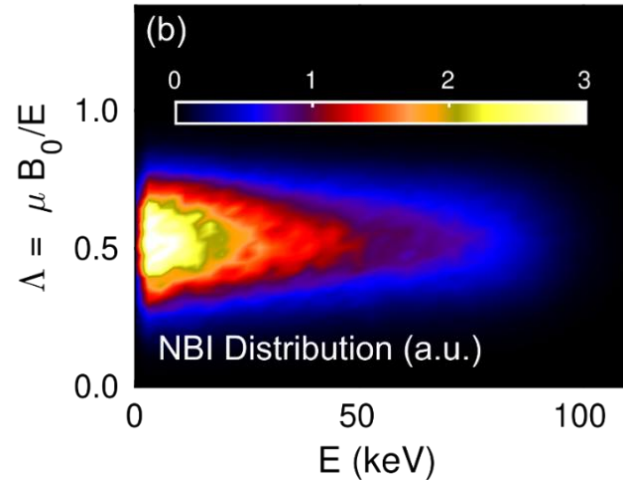
- Only when we have enough fast-ion population at the plasma edge, a significant impact on ELMs can be observed
- For core-localized distribution, energetic particle driven mode appears after ELM

Identifying core instability in simulations with core-localized distribution

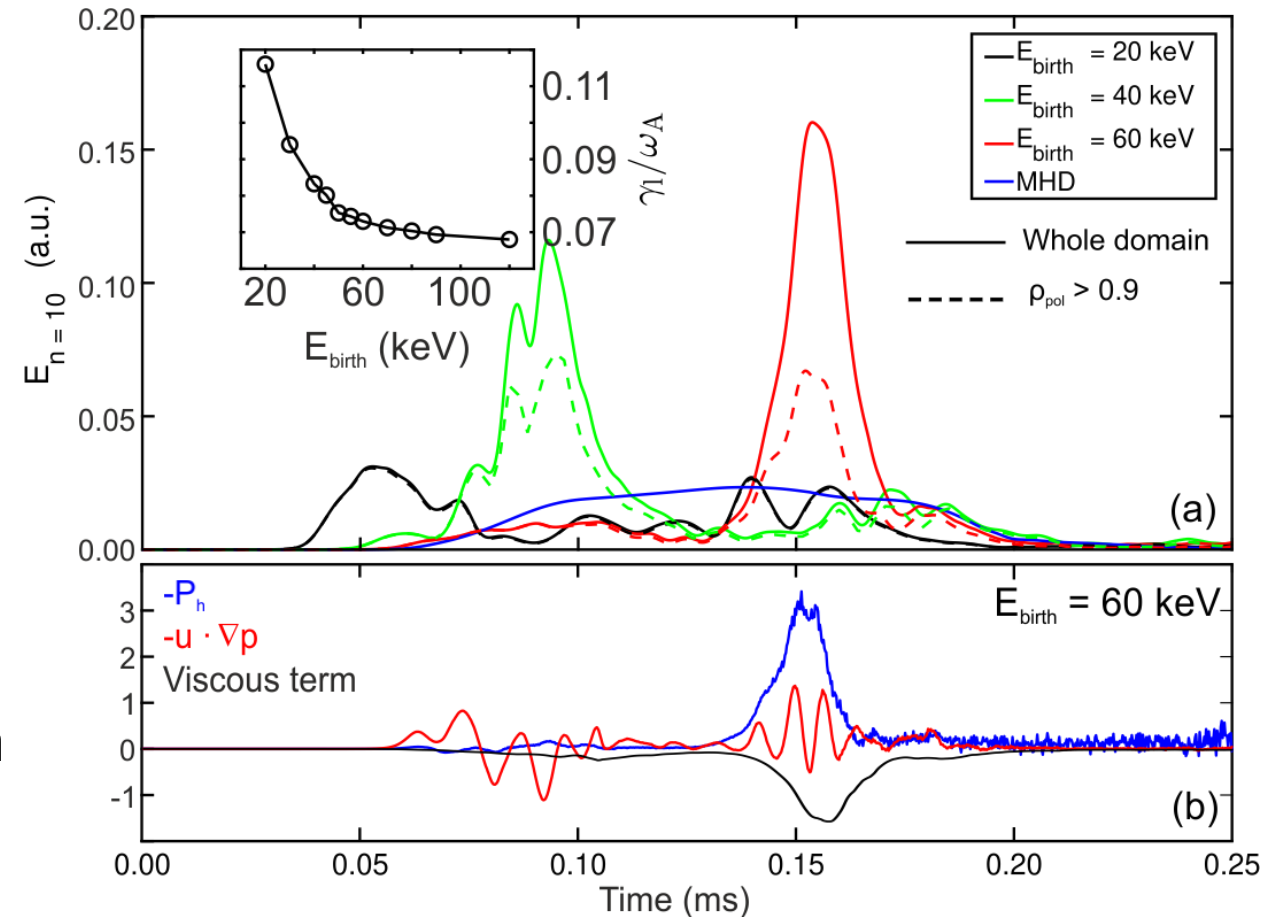
- The mode appearing after ELM is EPM
 - Does not fall in TAE gap
 - Discrete perturbation that crosses Alfvén continuum
- EPM saturates and perturbation extends radially inwards and outwards
- Motivates new study: **interplay between EP driven modes and ELMs**



Fast-ion birth energy impact on ELMs

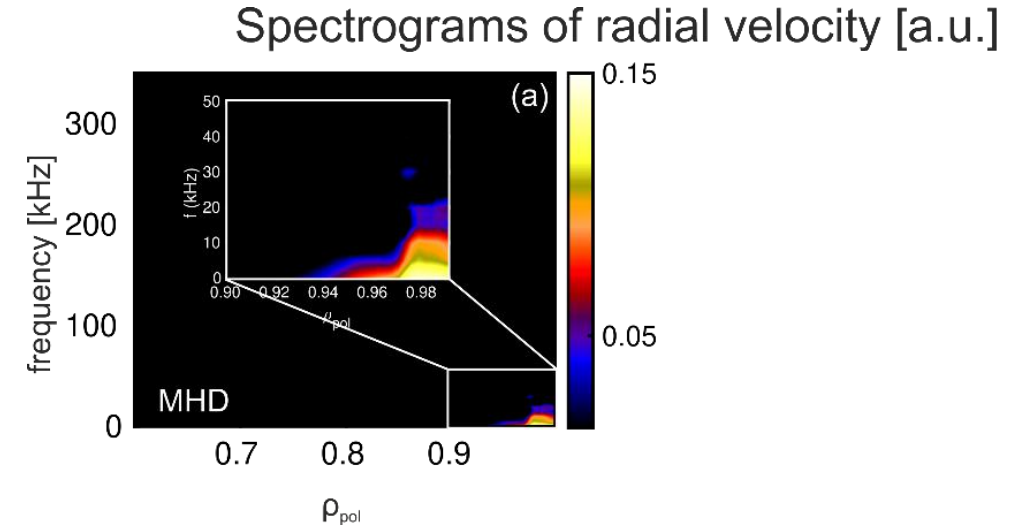


- γ decreases with E_{birth}
- ELM radial width is affected by fast-ion kinetic effects
- Hybrid simulations qualitatively reproduce abrupt crash (dominated by fast-ions) and pure MHD mild crash



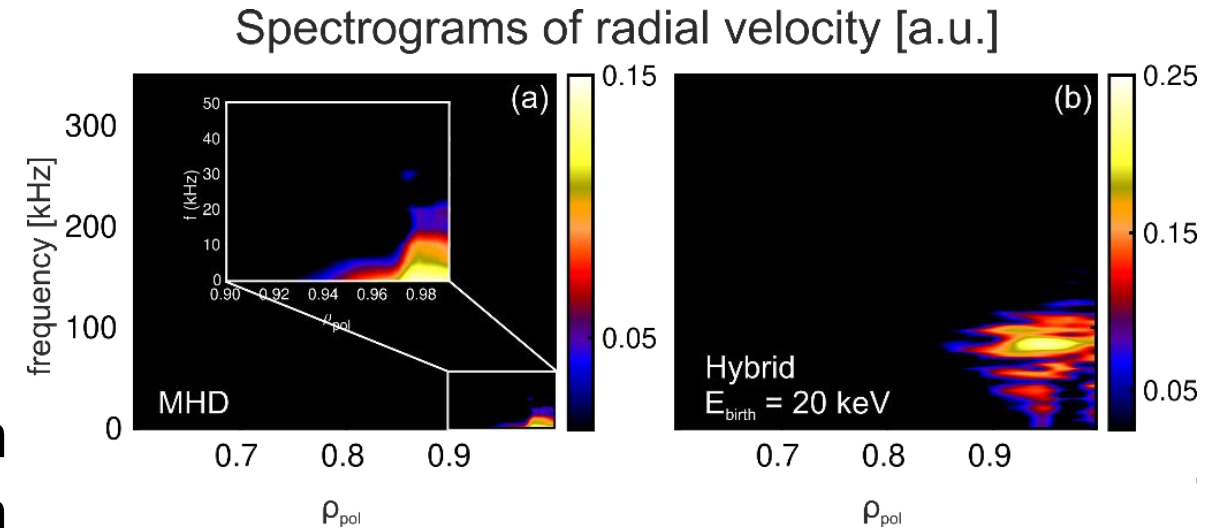
Spatial-frequency pattern of ELM in the presence of energetic particles

- In pure MHD simulation, $f < 20$ kHz



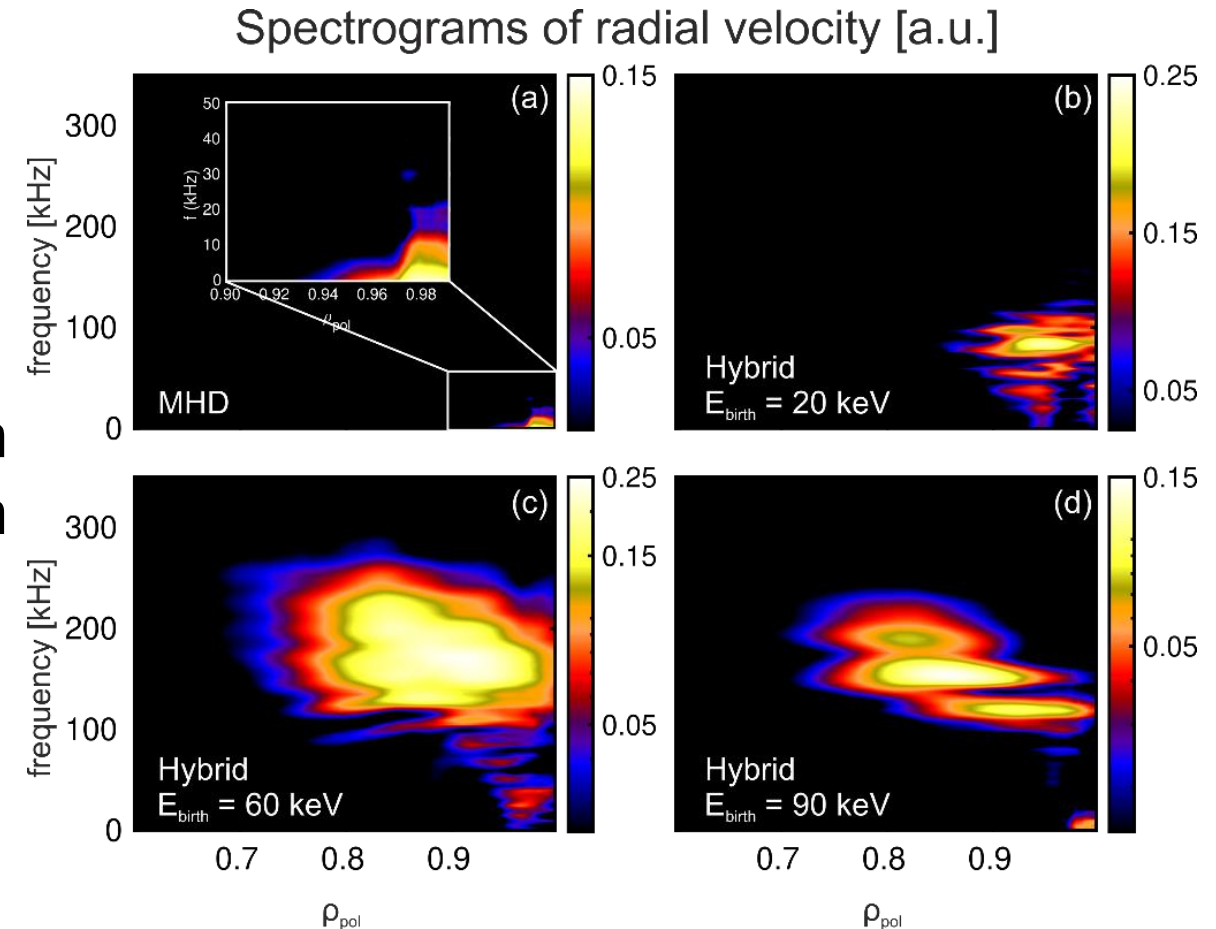
Spatial-frequency pattern of ELM in the presence of energetic particles

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- In hybrid simulations, ELM crash frequency pattern depends on fast-ion energy with dominant $f > 100$ kHz

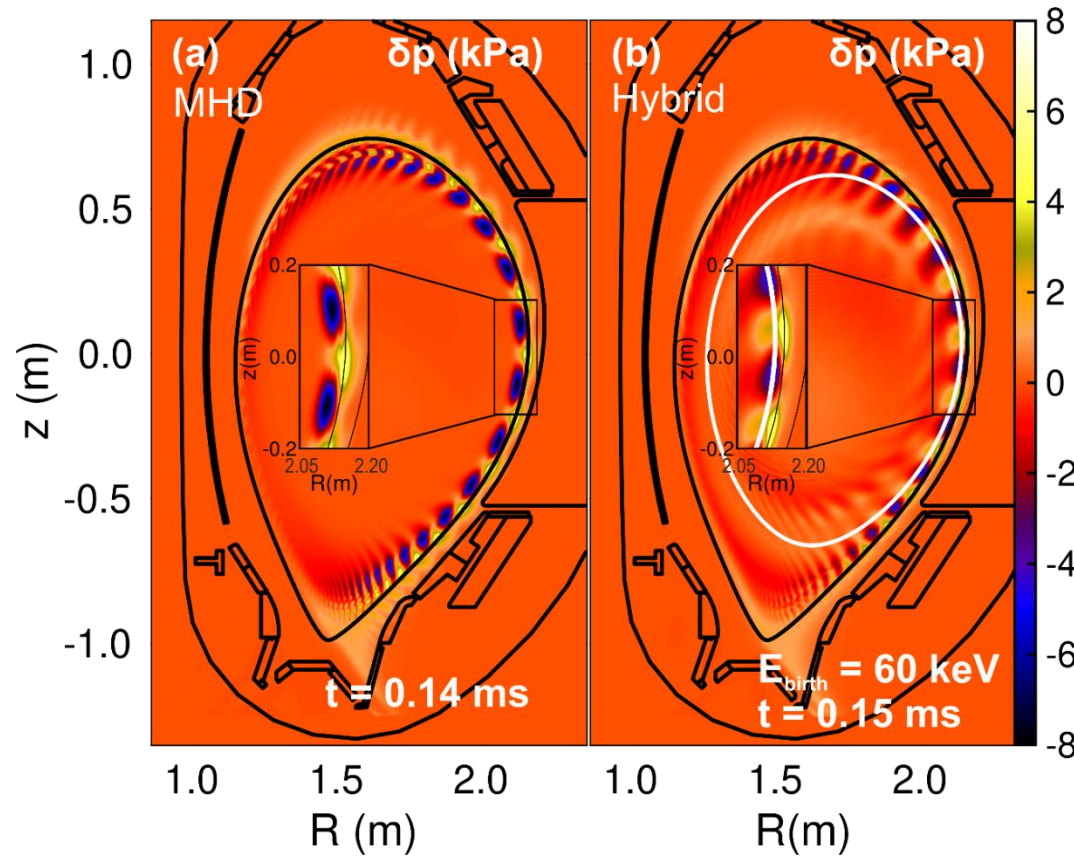


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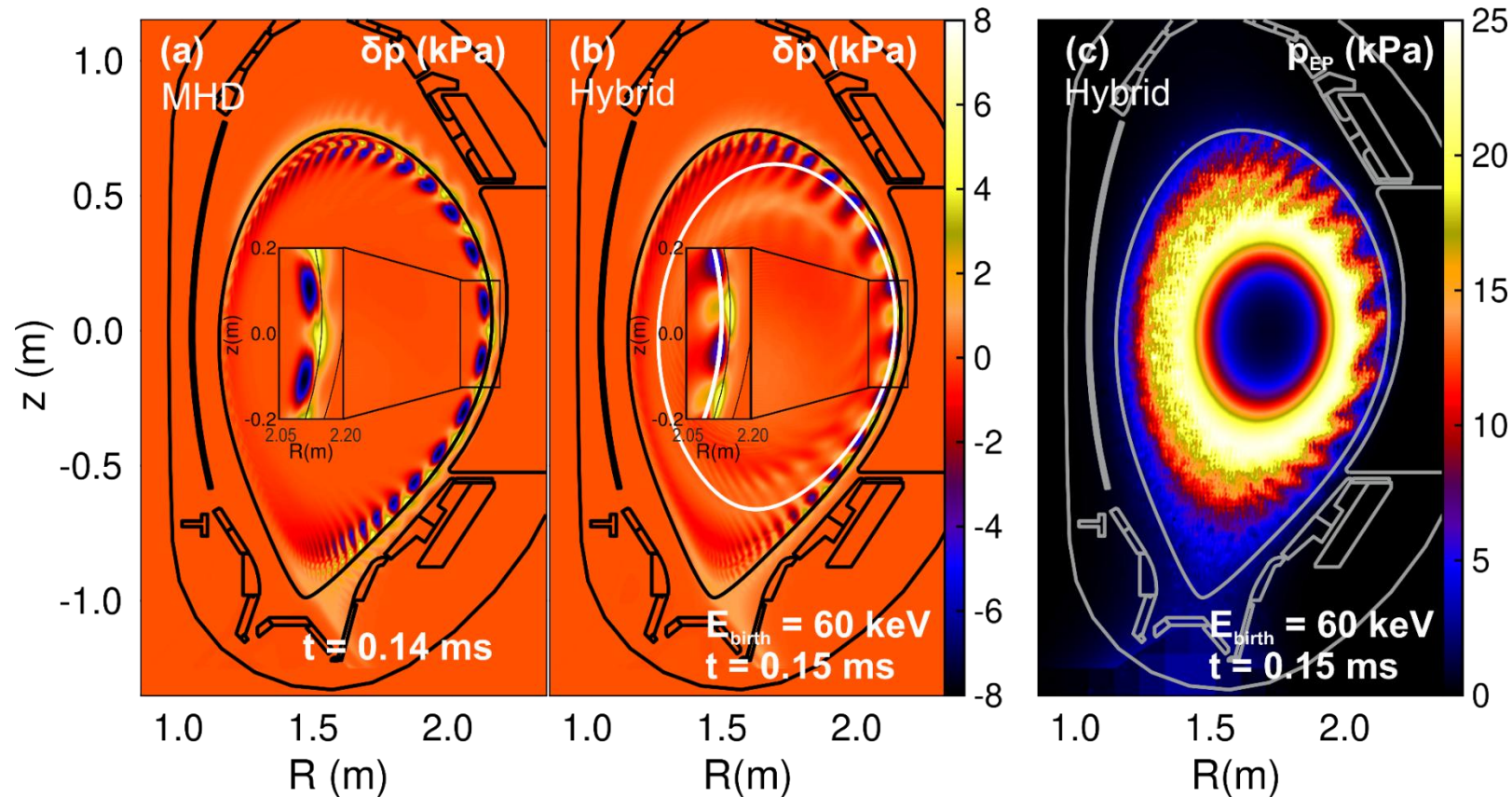


Fast-ion kinetic effects alter ballooning mode structure



- Fast-ion kinetic effects modify the shearing of the ballooning mode

ELMs cause significant fast-ion redistribution



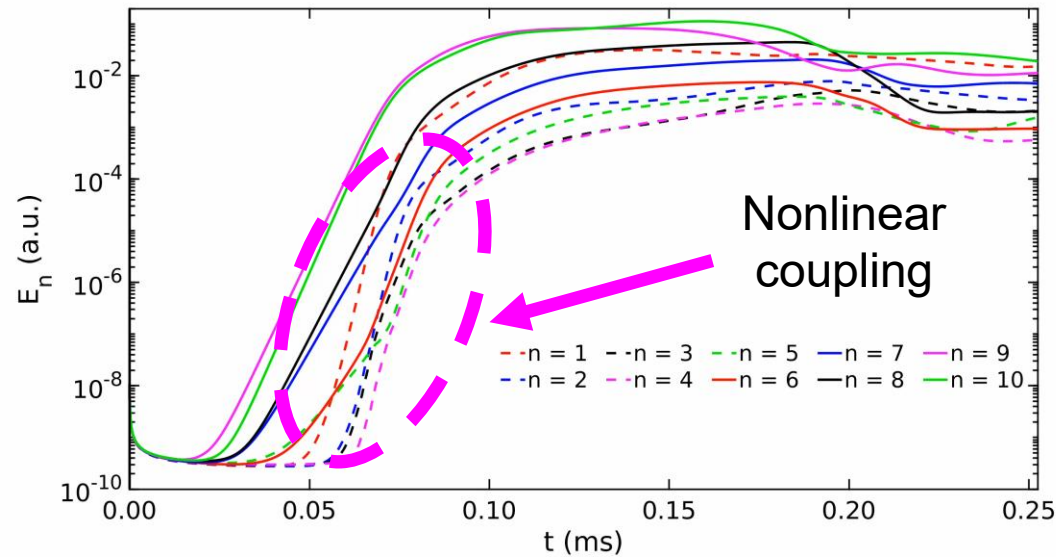
- Fast-ion kinetic effects modify the shearing of the ballooning mode
- ELMs cause fast-ion redistribution with ballooning, finger like structure with significant radial extension

The energy release in fusion is huge!

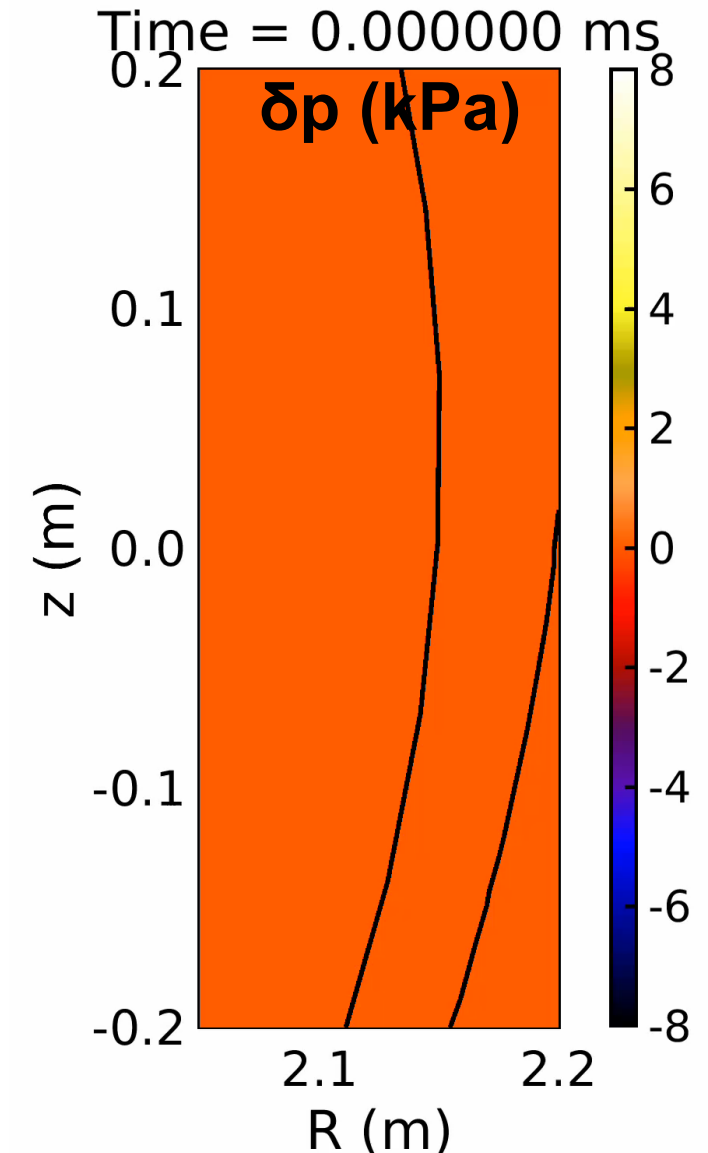
	combustion energy		
	$\left(\frac{\text{kcal}}{\text{mol}}\right)$	$\left(\frac{\text{kcal}}{\text{g}}\right)$	$\left(\frac{\text{kJ}}{\text{g}}\right)$
hydrogen (H_2)	68	34.0	142
methane (CH_4)	213	13.3	56
gasoline (C_8H_{18})	1303	11.5	48

- 2 tons of carbon produces $\sim 10^3$ kJ to power household's electrical appliances
- The energy release in D-T reaction is $\sim 10^{11}$ kJ / g
- It is huge compared to typical combustion reactions!

How does an ELM look like in a MHD simulation?



- Initial growth of high- n ballooning modes
- Nonlinear coupling between modes
- Saturation $\rightarrow$ particle release $\rightarrow$ pressure flattening

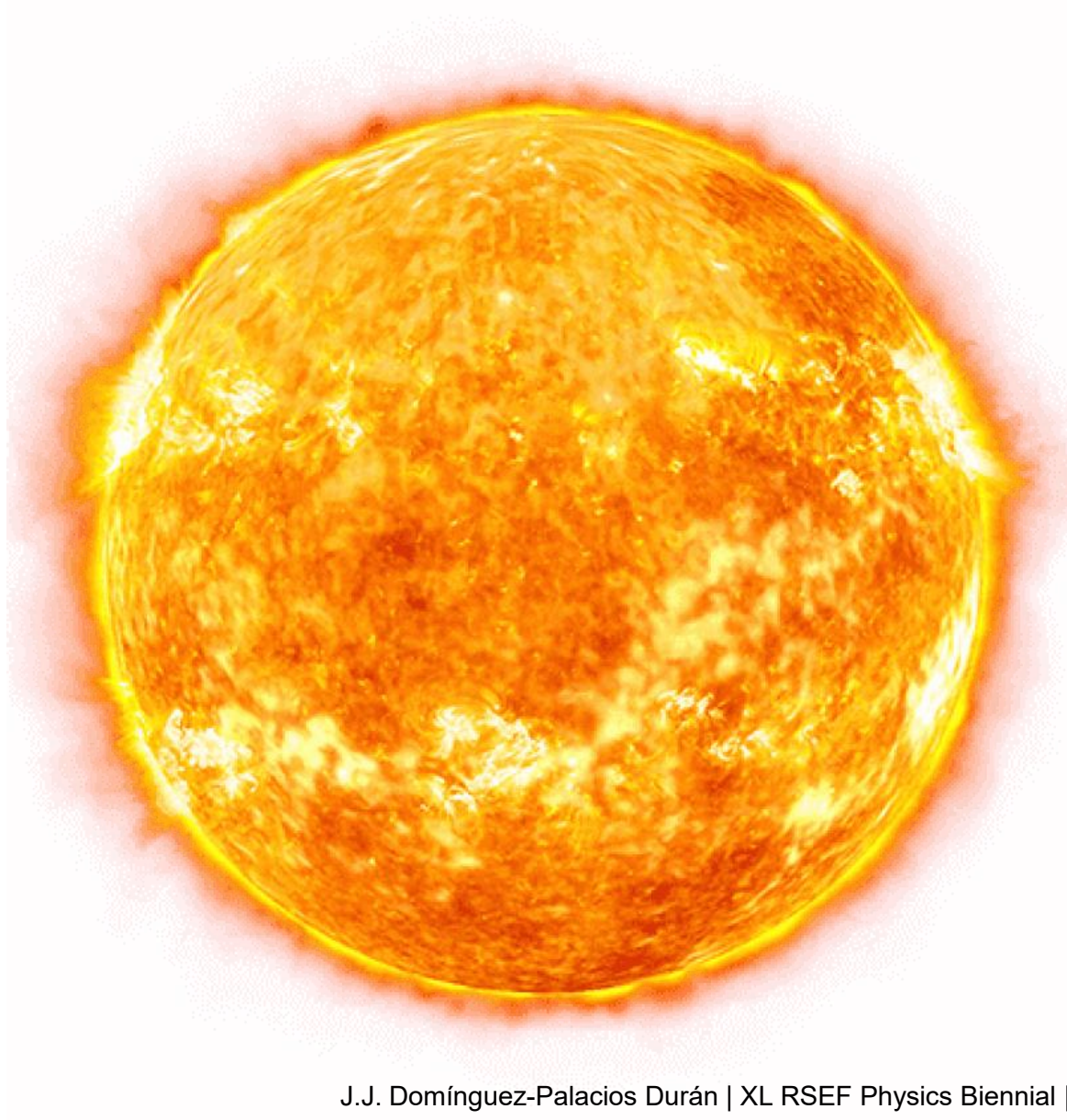


The world energy demand relies mostly on fossil fuels



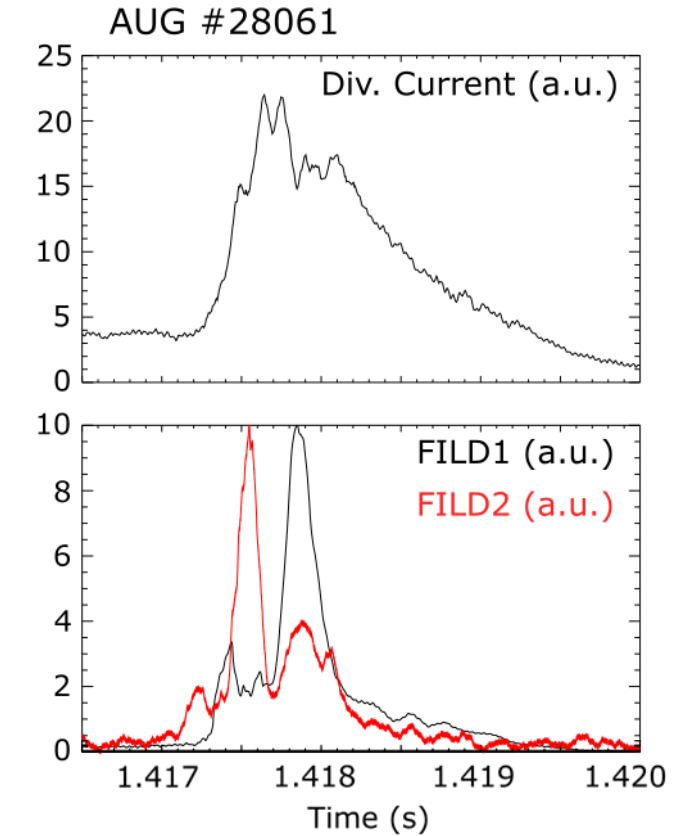
Statistical Review of World Energy 2025 by Energy Institute

Nuclear fusion: the energy source of stars



Experimental observations suggest interaction between ELMs and fast-ions

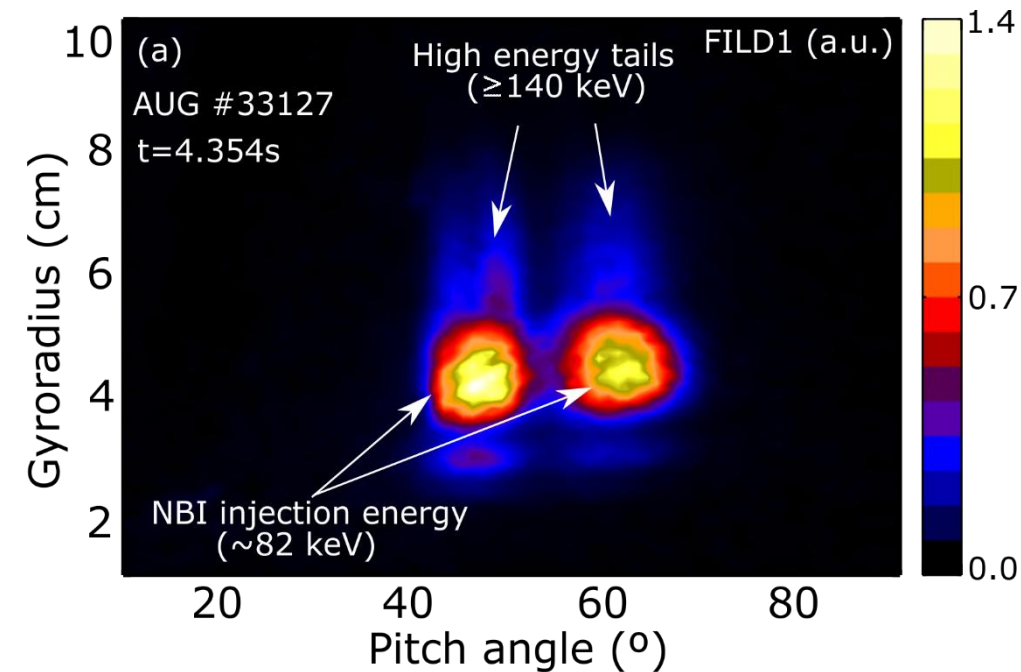
- Fast-ion losses measured during ELM crashes
 - FILD1 and FILD 2 located at same θ and different φ , both measuring FI losses



*J. Galdon-Quiroga *et al.*, 2018, Phys. Rev. Lett. **121**, 025002

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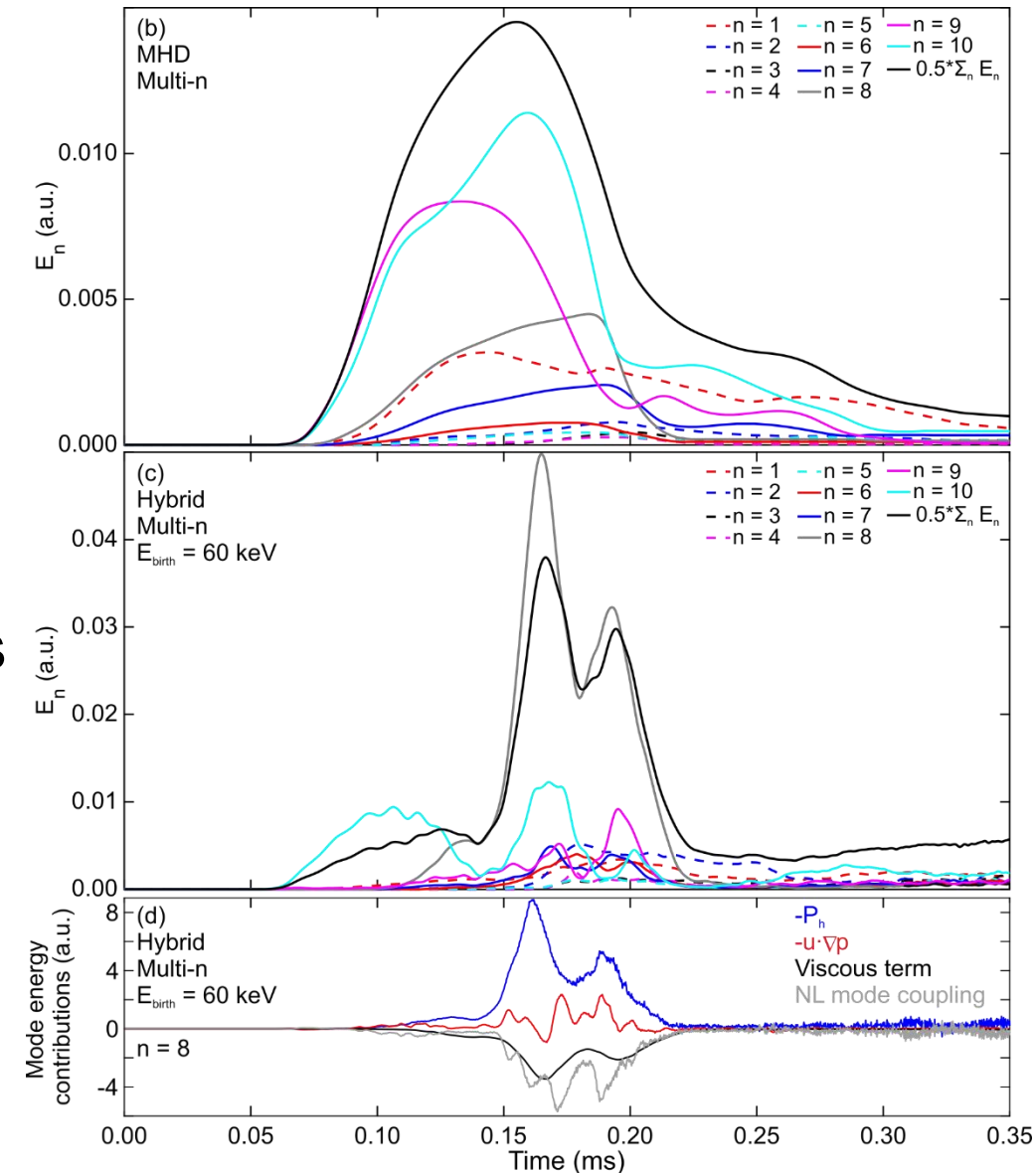
- Fast-ion losses measured during ELM crashes
 - FILD1 and FILD 2 located at same θ and different φ , both measuring FI losses
- Beam-ion acceleration during ELM crashes



*J. Galdon-Quiroga *et al.*, 2018, Phys. Rev. Lett. **121**, 025002

Fast-ions strongly affect modes energy evolution in multi-n simulations

- MHD multi-n: $n = 9, 10$ are most dominant modes
- Fast-ions modify the most dominant mode
 - $n = 8$ most unstable with EP kinetic effects
- Peak mode energy larger in hybrid single-n simulation compared to $\sum_n E_n$

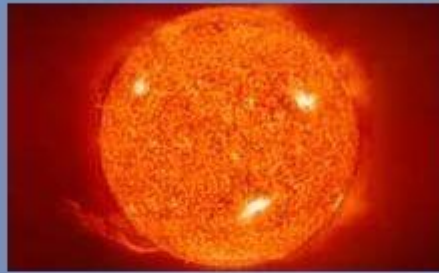


EJEMPLOS DE ESTADO DE PLASMA

Rayo



Sol



Luces de Neón



Aurora Boreal



Nebulosa



Arco de Soldadura

